

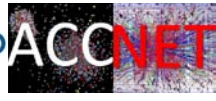
Machine Plans for Upgrades

“SLHC” type luminosities - issues and solutions

Frank Zimmermann
CMS Upgrade Week
29 April 2010



Input from 2001 LHC Upgrade Feasibility Study and from numerous CARE-HHH and EuCARD-AccNet workshops, LMC's, Chamonix2010, CERN MAC, etc.



Special thanks to R. Bailey, C. Bhat, O. Brüning, R. Calaga, H. Damerau, O. Dominguez, L. Evans, S. Fartoukh, R. Garoby, J.-P. Koutchouk, H. Maury Cuna, S. Myers, R. Ostojic, L. Rossi, F. Ruggiero, W. Scandale, G. Sterbini, L. Tavian, T. Taylor, E. Todesco, R. Tomas,...

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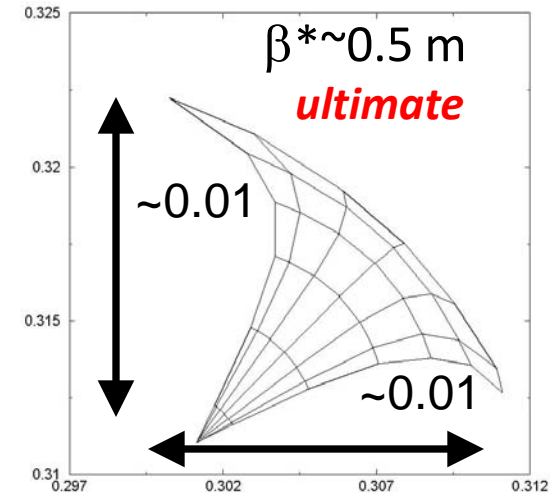
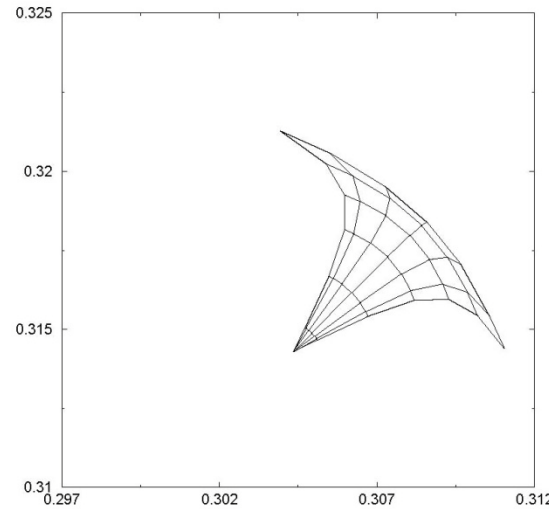
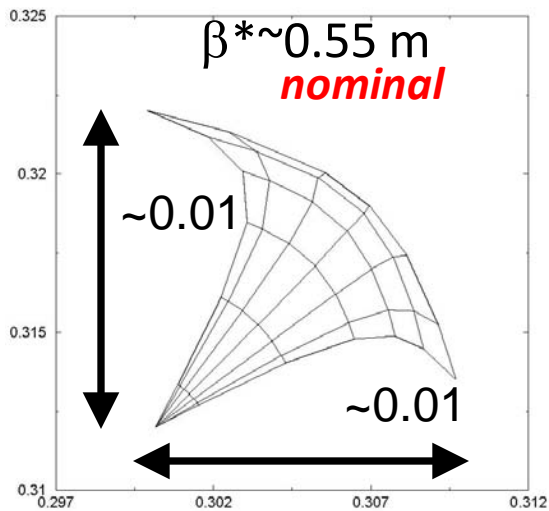
- **parameters available**
- **relationship, constraints and challenges**
- **possible parameter ranges**
- **different optimization strategies**
- **high luminosity scenarios**
- **upgrade time lines**

parameters

- β^* - IP beta function
- β_x^*/β_y^* - ratio of IP beta functions
- θ_c – (full) crossing angle
- ε_N – normalized transverse emittance
- N_b – bunch intensity
- n_b – number of bunches ($\rightarrow s_b$ - bunch spacing)
- longitudinal bunch profile (“flat” vs “Gaussian”)
- number of collision points (IP’s)
- T_{ta} – turn-around time

#IP's : the original plan – “phase 0”

J.Gareyte, F. Ruggiero *et al*, e.g. LHC'99 workshop, LHC Project Report 626



nominal tune footprint
up to 6σ with **4 IPs & nom.**
intensity $N_b = 1.15 \times 10^{11}$

$$L = 10^{34} \text{ cm}^{-2}\text{s}^{-1}$$

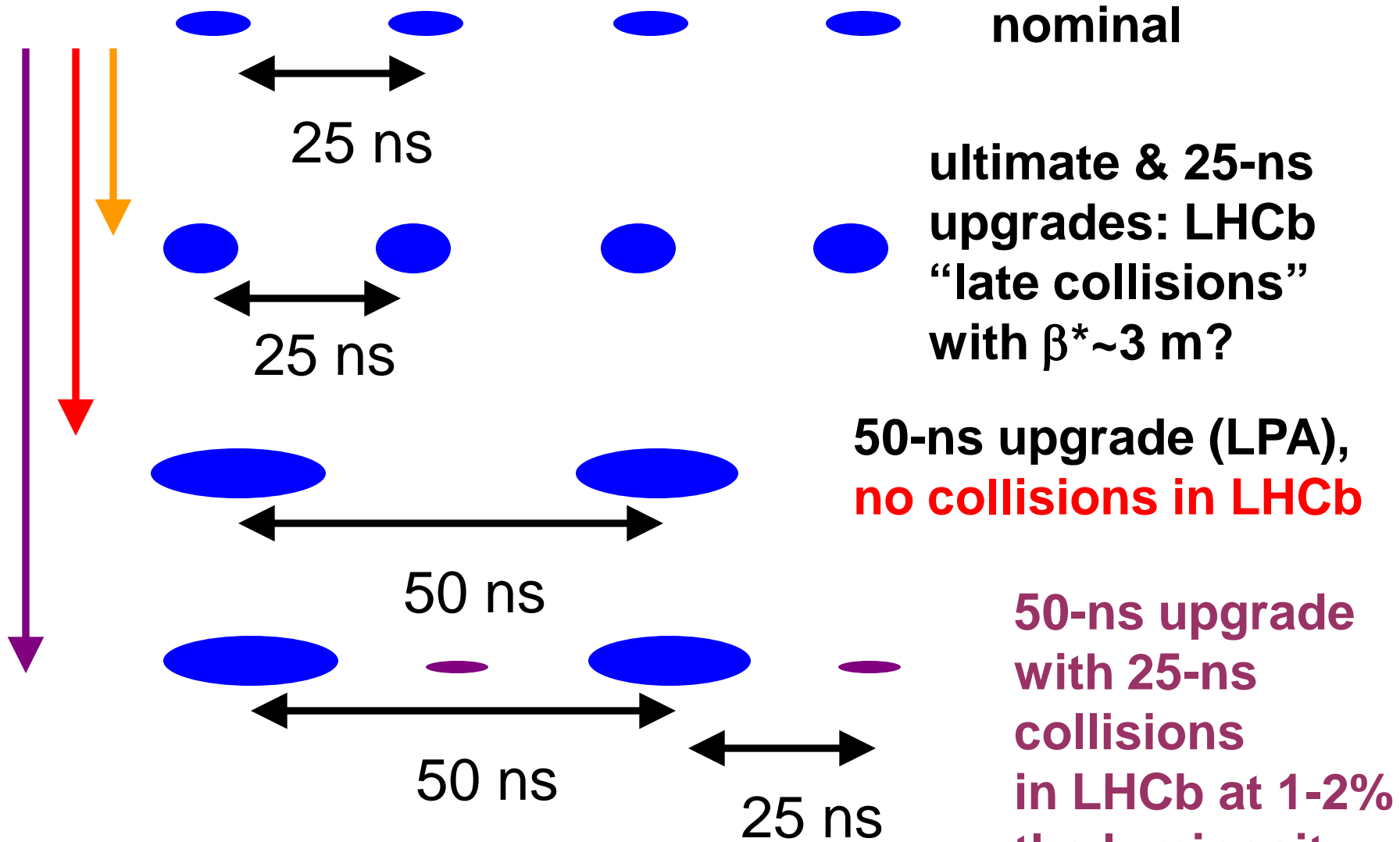
tune footprint up to 6σ
with nominal intensity
and **2 IPs**

tune footprint up to 6σ
with **2 IPs at ultimate**
intensity $N_b = 1.7 \times 10^{11}$

$$L = 2.3 \times 10^{34} \text{ cm}^{-2}\text{s}^{-1}$$

“going from 4 to 2 IPs ATLAS & CMS luminosity can be increased by factor 2.3 - further, increasing crossing angle to $340 \mu\text{rad}$, bunch length (x2), & bunch charge to $N_b = 2.6 \times 10^{11}$ would yield $L = 3.6 \times 10^{34} \text{ cm}^{-2}\text{s}^{-1}$ [$\beta^* = 0.5$ m]”

what about LHCb? – bunch patterns



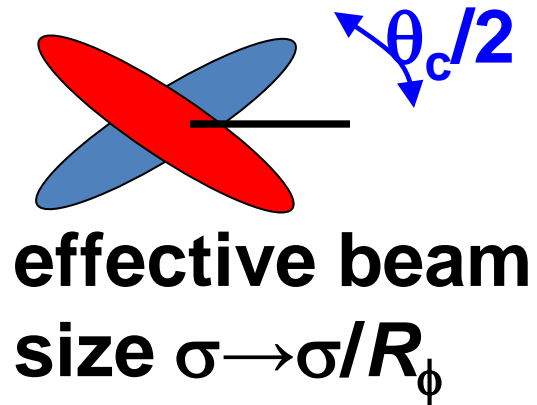
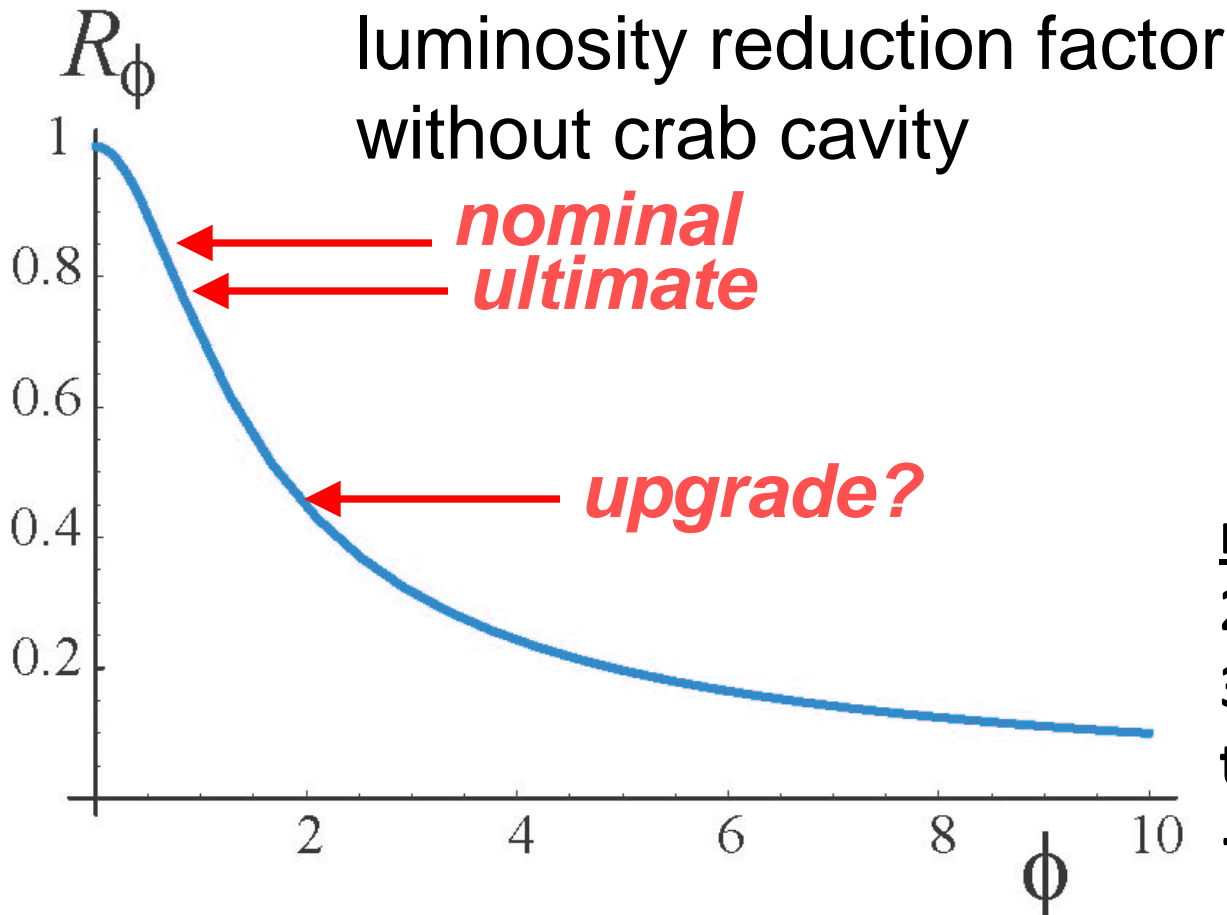
50 ns: much reduced e- cloud! LHCb transparent

constraints

- **total beam-beam tune shift ≤ 0.01**
 - SPS p-pbar experience
- long-range beam-beam \rightarrow **crossing angle $\geq 9\sigma$**
- **arc cooling capacity**
 - global & local limitations, cooling shares with IR
 - heat load from SR, image currents, & e-cloud
- IR layout & optics $\rightarrow \beta^*$
- **event pile up** in the detectors (≤ 300 , ≤ 200 ?)
- **luminosity lifetime** ($\geq 2\text{h}$? $\geq 5\text{h}$?)

constraint - crossing angle

$$R_\phi = \frac{1}{\sqrt{1 + \phi^2}}; \quad \phi \equiv \frac{\theta_c \sigma_z}{2\sigma_x^*} \quad \text{“Piwinski angle”}$$



range - $f(\text{triplet}, \beta^*)$:
285 μrad (nominal)
315 μrad (ultimate)
till $\sim 410 \mu\text{rad}$ “phase I”
 $\rightarrow 500 \mu\text{rad}$ “phase II”?

beam-beam tune shift, ϕ & luminosity

$$\Delta Q_{bb} = \frac{N_b}{\gamma \varepsilon} \frac{r_p}{2\pi} \frac{1}{\sqrt{1 + \phi_{piw}^2}} \frac{1}{F_{profile}}$$

total b-b tune shift
for two IP's with
alternating crossing

$$L = \frac{1}{4\pi} f_{rev} n_b \gamma \frac{1}{\beta^* (\gamma \varepsilon)} N_b^2 \frac{1}{\sqrt{1 + \phi_{piw}^2}}$$

at the b-b limit, larger Piwinski angle &/or larger emittance increase luminosity!

$$= \frac{\pi}{r_p^2} f_{rev} n_b \gamma \frac{(\gamma \varepsilon)}{\beta^*} \Delta Q_{bb}^2 F_{profile}^2 \sqrt{1 + \phi_{piw}^2}$$

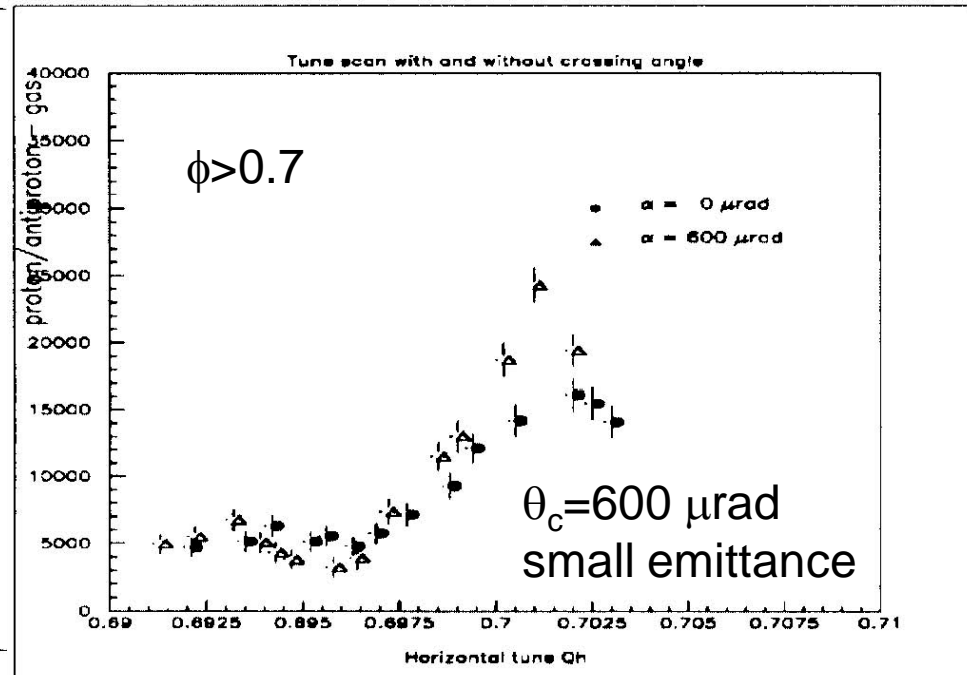
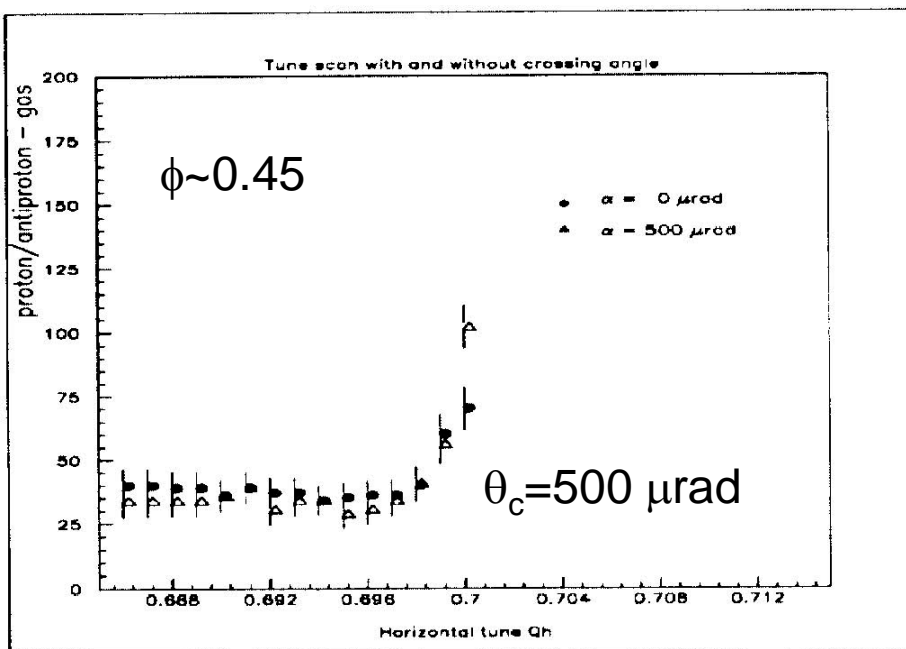
optimization strategies:

- 1) increase N_b with ε (e.g. controlled ε blow up at top energy)
- 2) increase N_b with $1/R_\phi$ & “flat” bunch $F_{profile} \sim 1.4$ (“LPA”)
- 3) vary ε as $1/R_\phi$ (“small emittance”)
- 4) set $1/R_\phi = 1$ at IP and minimize β^* (e.g. crab crossing)

beam-beam limit – θ_c dependence?

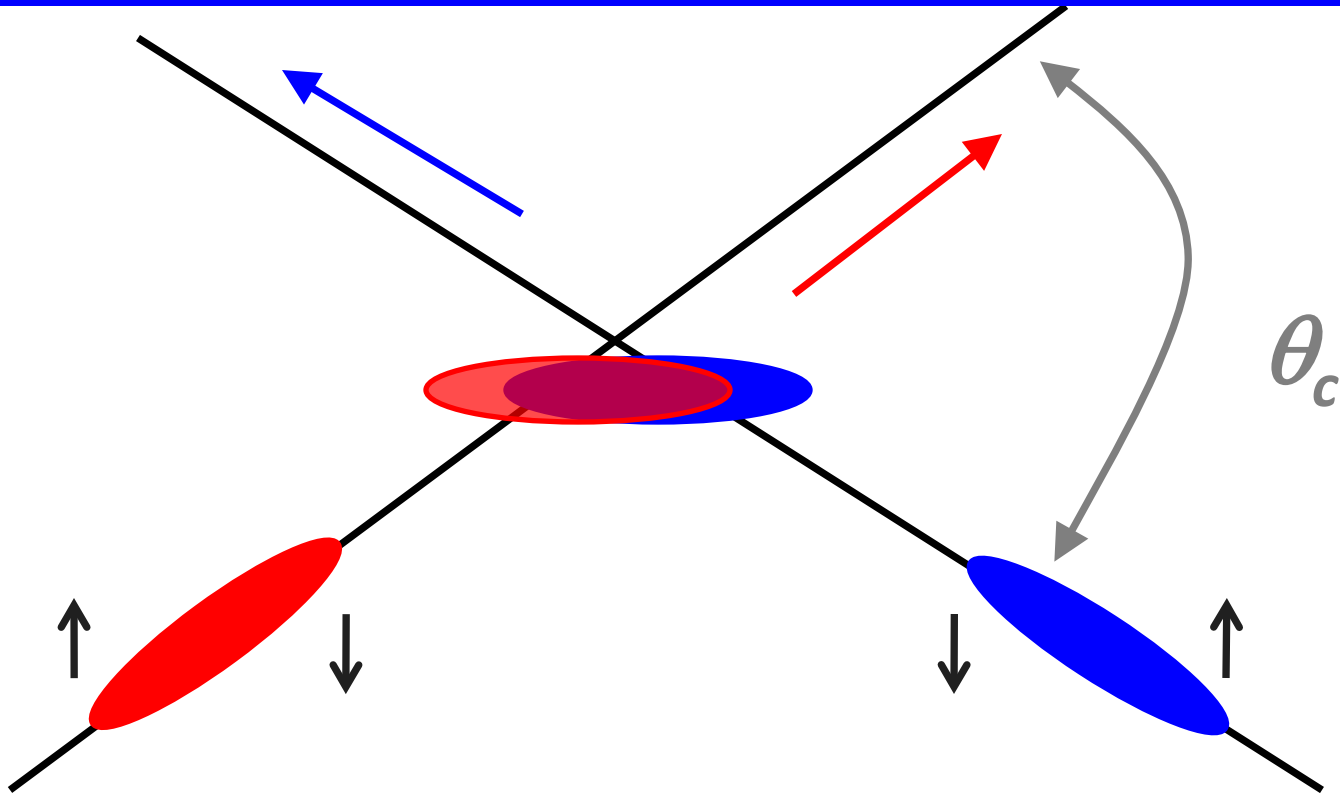
in lepton colliders crossing angle has reduced the beam-beam limit (DORIS-I, KEKB,...)

for hadrons, one historical experiment at the SPS
K. Cornelis, W. Herr, M. Meddahi, PAC91 San Francisco



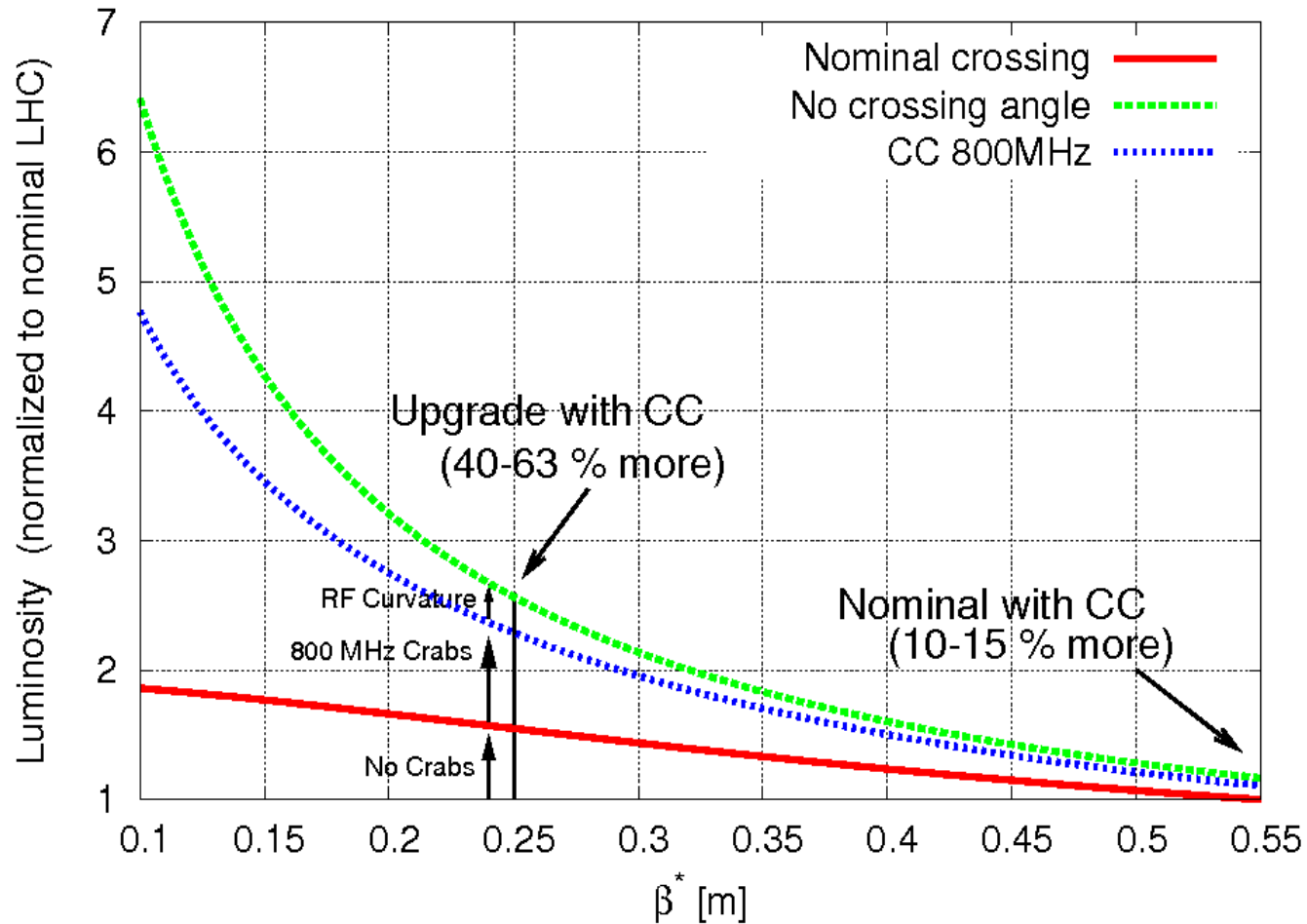
(almost) no additional beam-beam effect, but ϕ was much smaller than considered for SLHC

crab crossing



- RF crab cavity deflects head and tail in opposite direction so that collision is effectively “head on” for luminosity and tune shift
 - bunch centroids still cross at an angle (easy separation)
 - 1st proposed in 1988, in operation at KEKB since 2007
- advantages: higher geometric luminosity, easy leveling, potentially higher beam-beam tune shift**

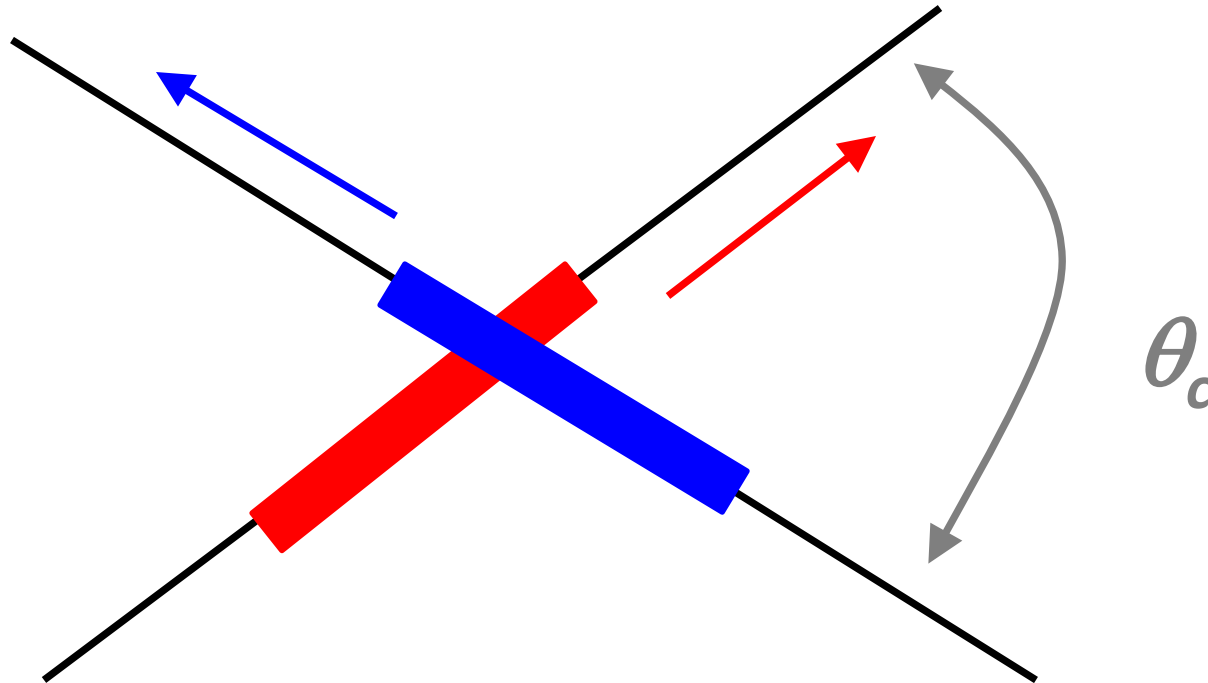
peak luminosity gain w crab cavities



R. Calaga

crab cavities are important & effective for small β^*

large Piwinski angle – “LPA”



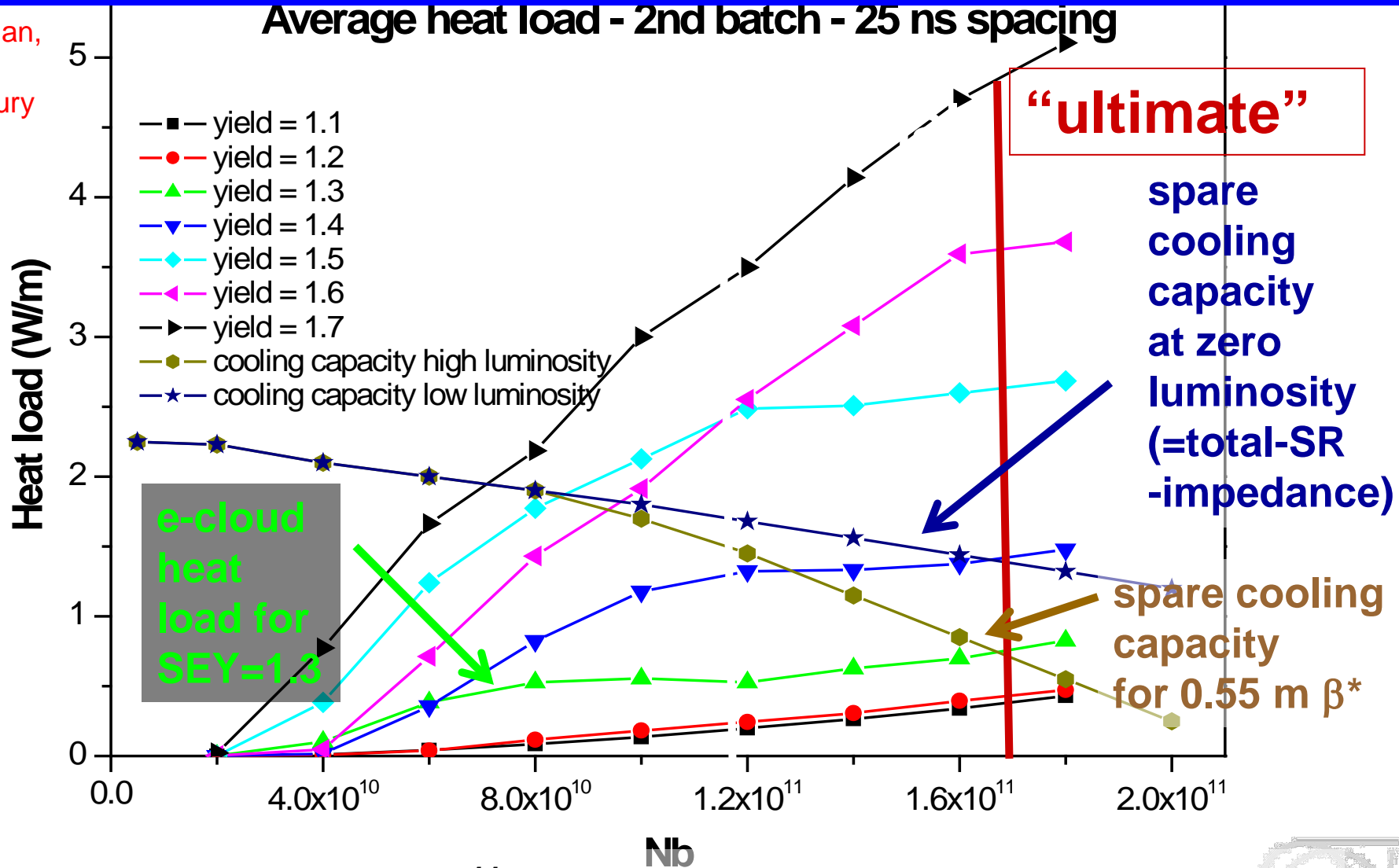
- 1) large Piwinski angle $\theta_c \sigma_z \gg 2 \sigma_x^*$
 - 2) longitudinally flat profile
- reduced tune shift, higher bunch charge
(& 50 ns spacing for e-cloud)**

constraints - N_b range

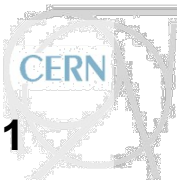
- **beam-beam tune shift** of “head-on” collision
 - ✓ is the limit for crab crossing;
 - ✓ going beyond ultimate N_b requires large Piwinski angle or large emittance;
 - ✓ even larger crossing angle than for LR-BB may be needed
- **arc cooling capacity**
- **injectors, collimation**, machine protection,...

cooling & e- heat for 25 ns spacing

L. Tavian,
2005
H. Maury
Cuna,
2009

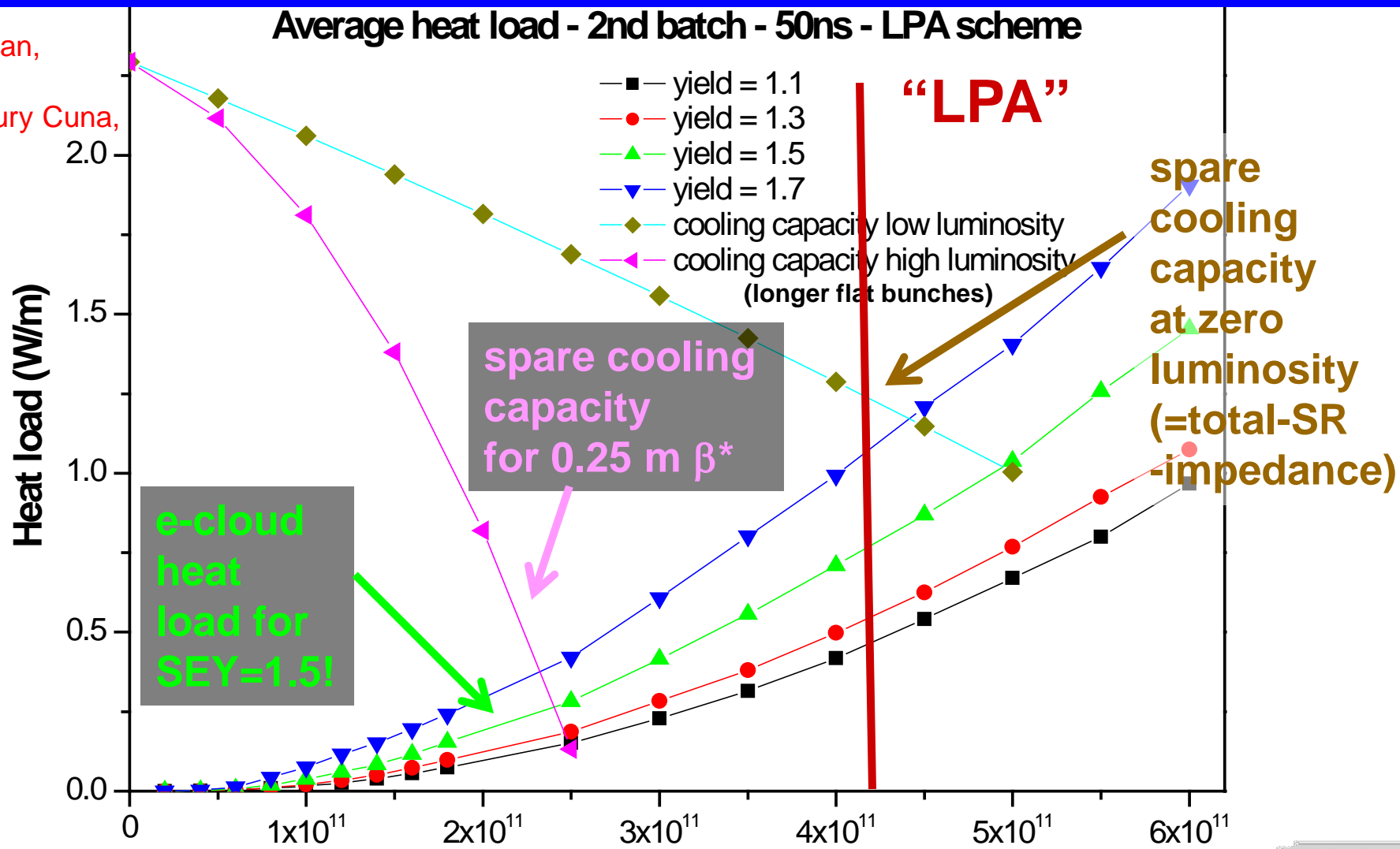


going above $N_b=1.7 \times 10^{11}$ & ultimate luminosity requires dedicated IR cryo plants; limit then becomes $N_b \sim 2.3 \times 10^{11}$

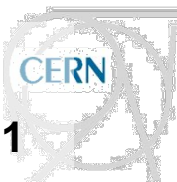


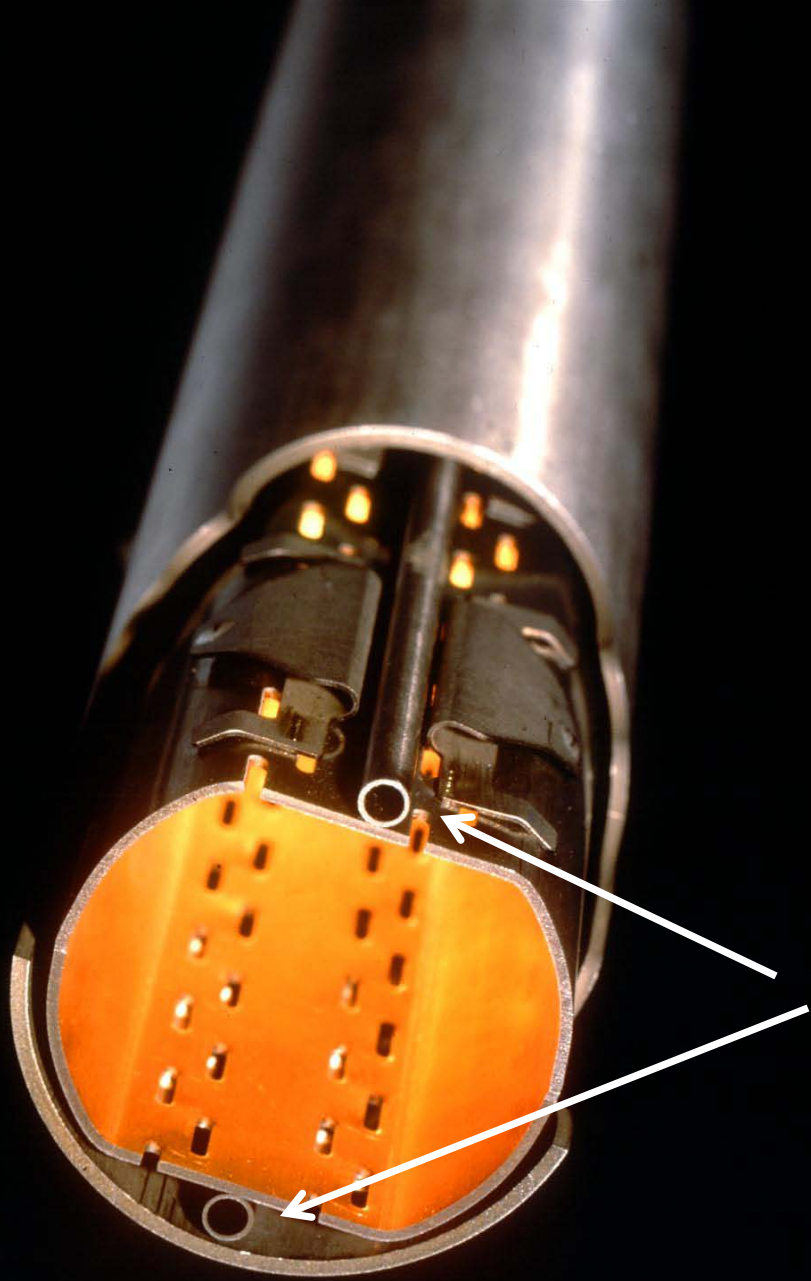
cooling & e- heat for 50 ns spacing

L. Taviani,
2005
H. Maury Cuna,
2009



going above $N_b = 2.3 \times 10^{11}$ & ultimate luminosity requires dedicated IR cryo plants; limit then becomes $N_b \sim 5.0 \times 10^{11}$



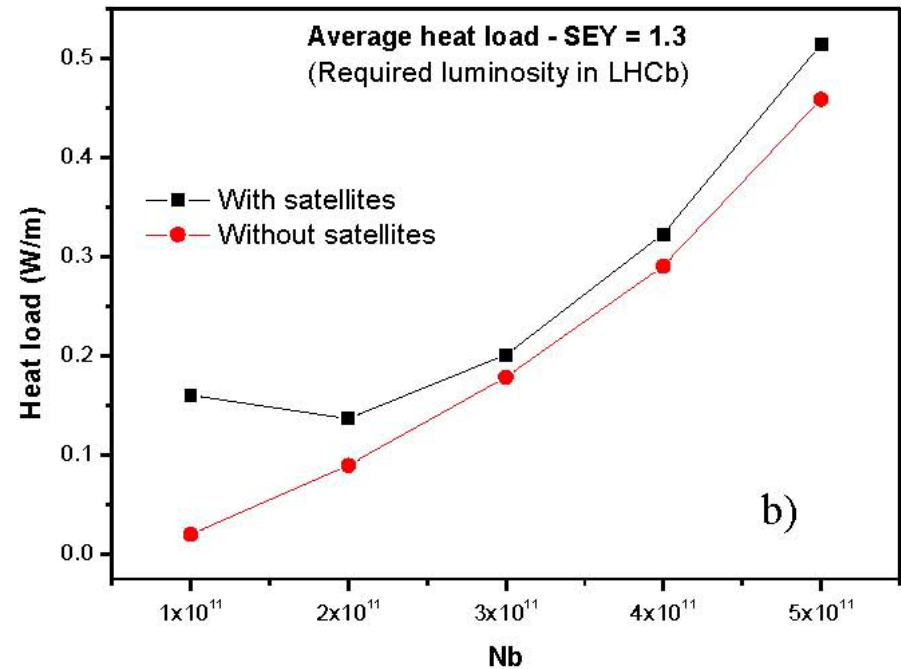
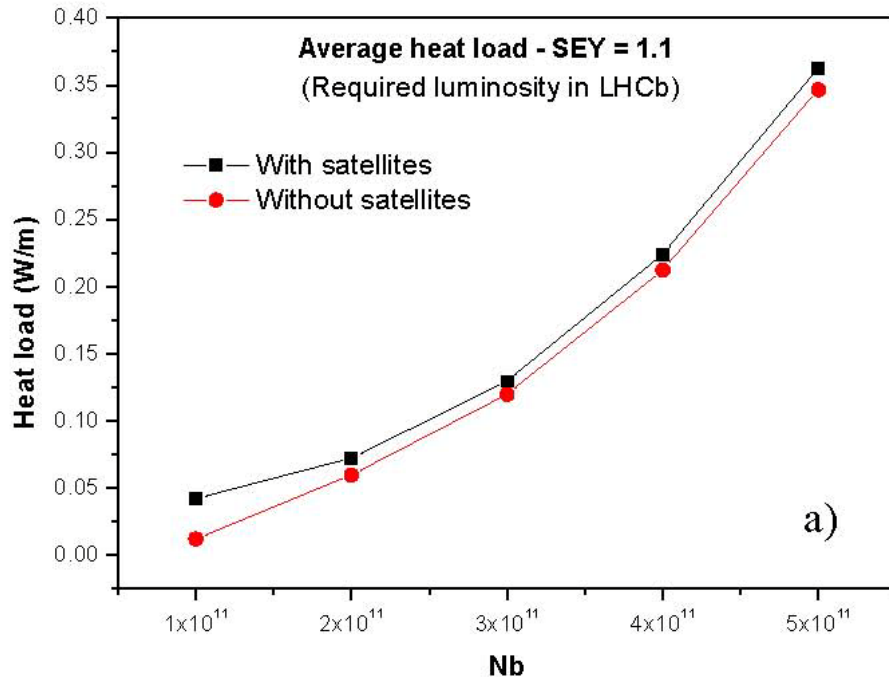


with separate
IR cryoplane,

the local arc
cooling capacity is
limited by hydraulic
impedance (i.e. the
diameter) of the
cooling capillaries
on the arc
beam screens

e- heat with LHCb satellite: OK

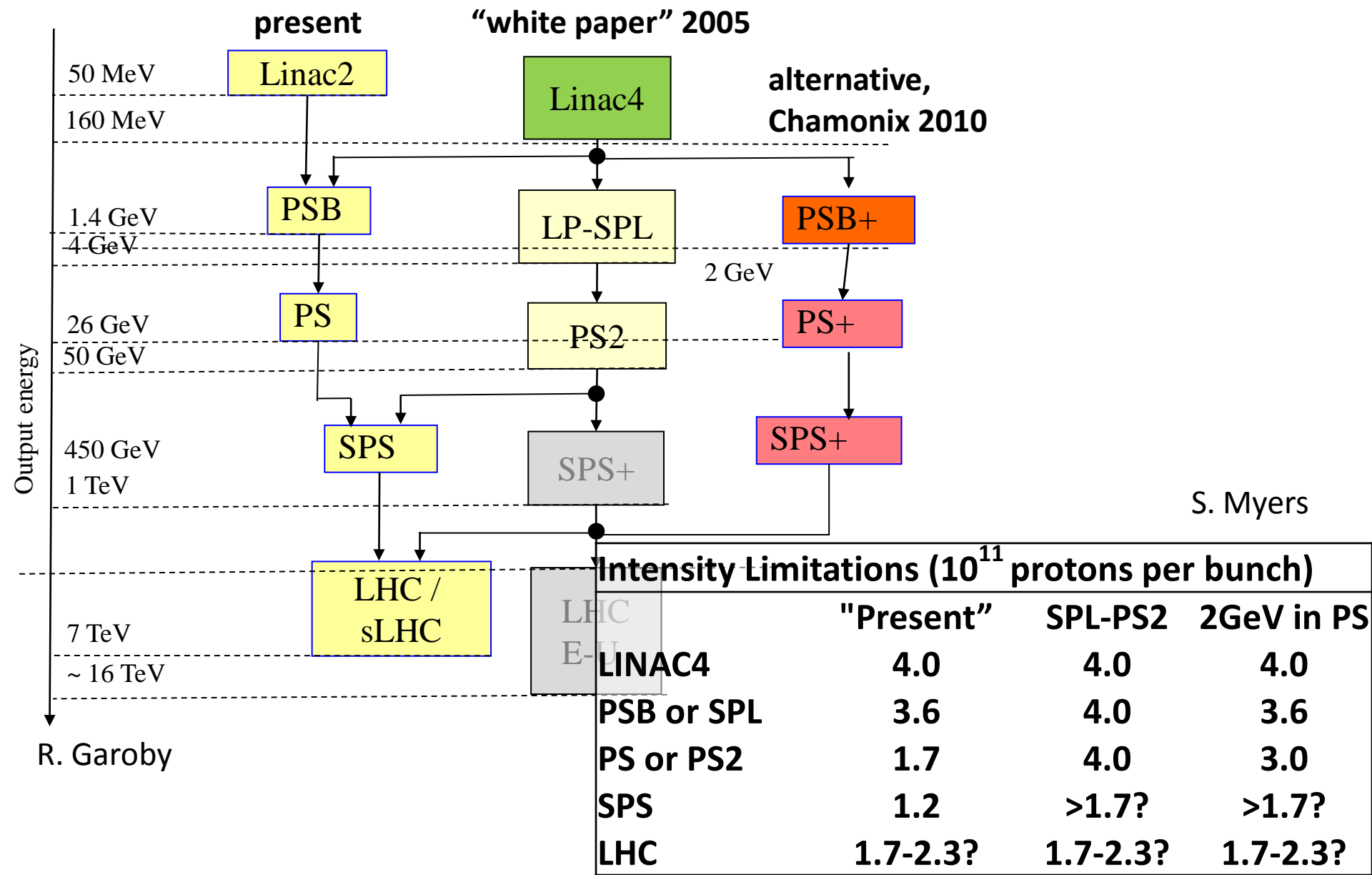
H. Maury Cuna, 2009



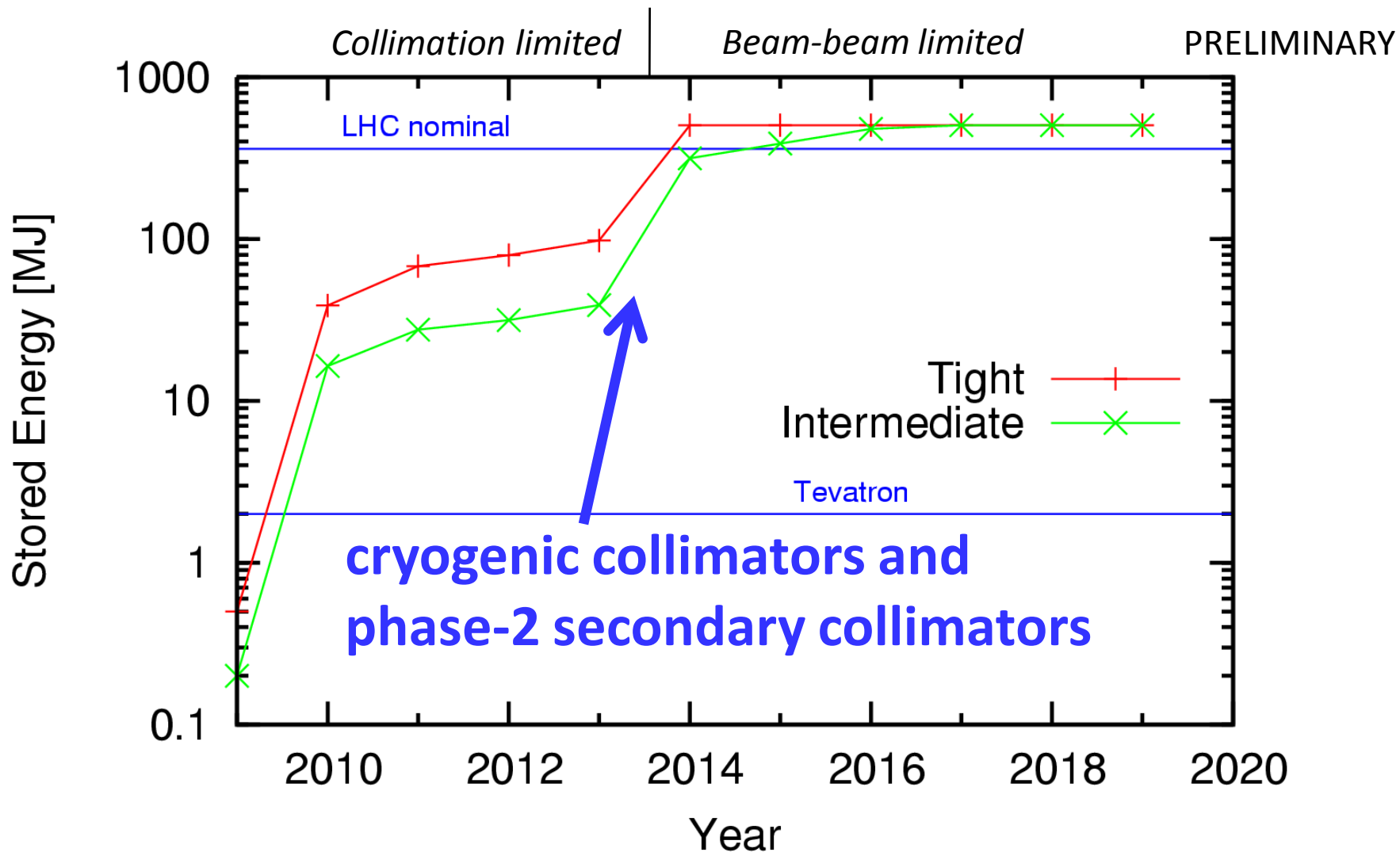
satellite intensity is varied as the inverse of main-bunch intensity to yield target luminosity of $2 \times 10^{33} \text{ cm}^{-2} \text{ s}^{-1}$ in (S)LHCb

“LHCb satellite” has small effect on 50-ns heat load

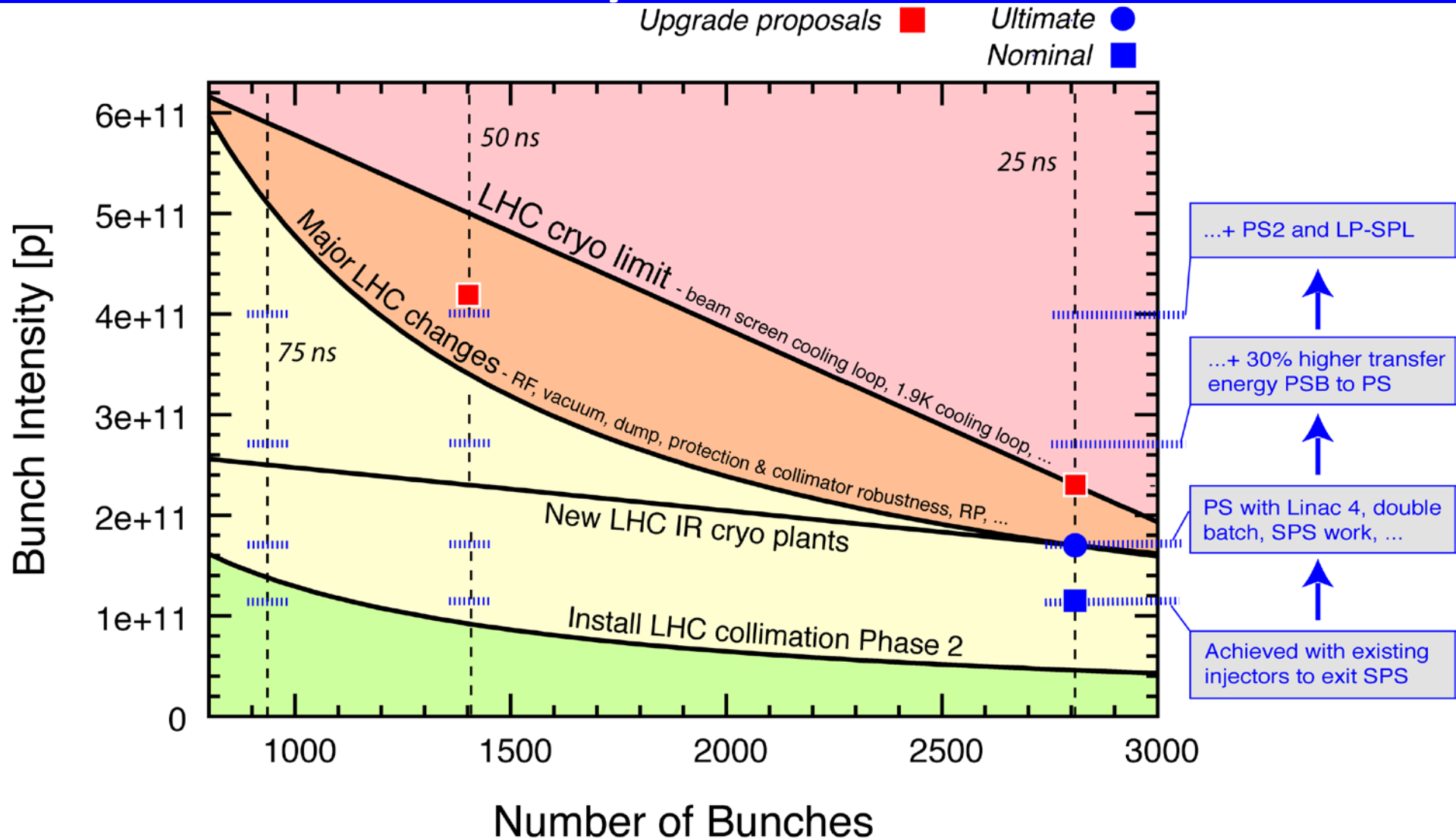
injector upgrade & intensity limits



collimation intensity limits



LHC intensity limits - schematic



constraint – β^* range

0.55 m nominal

0.50 m ultimate

0.40 m

0.30 m

0.25 m

0.22 m

...

0.14 m

...

0.07 m?!

IR “phase I” ,

larger aperture NbTi quad’s +...

IR “phase II”

Nb₃Sn quad’s + ...

novel Q’ correction schemes

constraint – pile up

bunch collision rate

= #bunches/beam x revolution frequency

#events per bunch crossing

= cross section x luminosity / bunch collision rate

nominal #events/crossing in the detector

= $6 \times 10^{-26} \text{ cm}^2 \cdot 10^{34} \text{ cm}^{-2} \text{ s}^{-1} / (32 \times 10^6 \text{ s}^{-1})$

= 19

inelastic cross section



e.g. 10 times higher luminosity at same #bunches

→ ~200 events per crossing (*detector upgrade!*)

luminosity decay & lifetime

fast decay of beam intensity and luminosity (few hours)
dominated by proton burn off

$$L(t) = \frac{\hat{L}}{\left(1 + t / \tau_{eff}\right)^2}$$

with

$$\tau_{eff} = \frac{N_b n_b}{n_{IP} \hat{L} \sigma_{tot}}$$

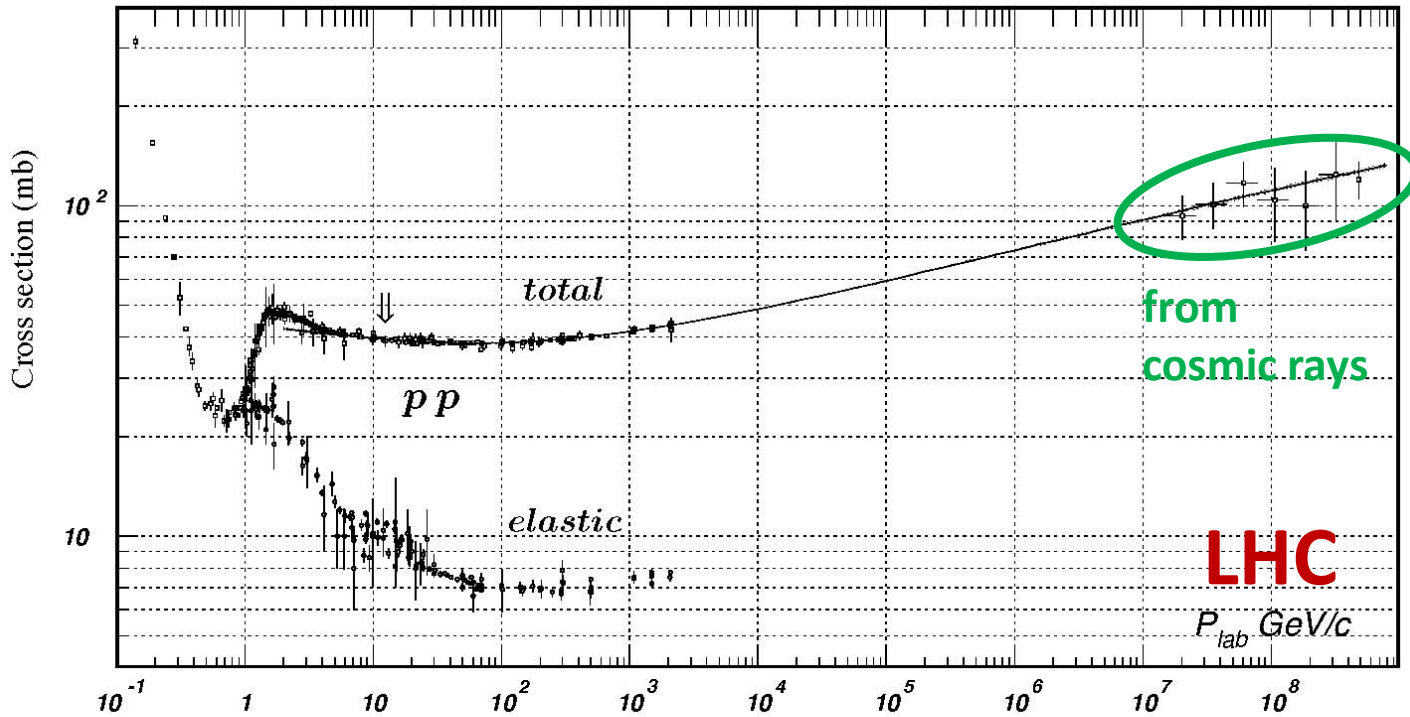
algebraic (\neq exponential) decay

$$\tau_{lumi} \propto \frac{\text{total beam intensity}}{\text{luminosity}}$$

for given luminosity, the luminosity lifetime
depends only on total beam current [w/o leveling]

cross sections

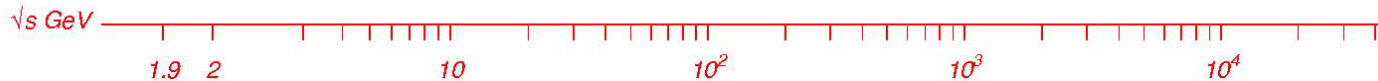
C. Amsler *et al.*, Physics Letters **B667**, 1 (2008)



$\sigma_{\text{tot}} \sim$
100 mbarn
 $\sim 10^{-25} \text{ cm}^2$

from
cosmic rays

$\sigma_{\text{inelastic}} \sim$
60 mbarn
 $6 \times 10^{-26} \text{ cm}^2$



total cross section for LHC c.m. energy from cosmic ray experiments

example scenarios

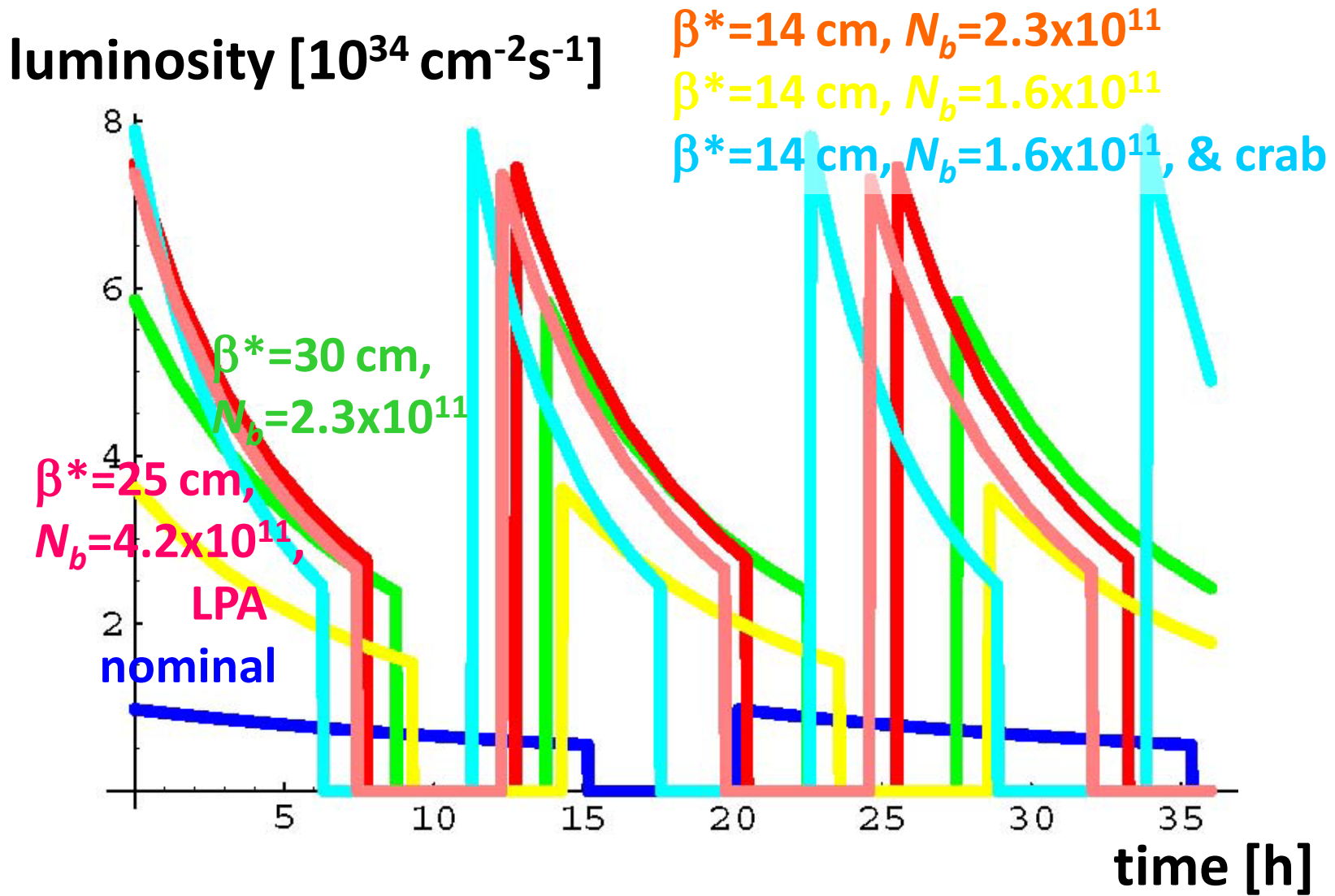
- (1) nominal, $N_b=1.15 \times 10^{11}$, $\beta^*=0.55$ m, $\theta_c=285$ μ rad
- (2) nominal*, $N_b=1.7 \times 10^{11}$, $\beta^*=0.55$ m, $\theta_c=285$ μ rad, 50 ns
- (3) ultimate, $N_b=1.7 \times 10^{11}$, $\beta^*=0.50$ m, $\theta_c=315$ μ rad
- (4) “phase I+”, $N_b=2.3 \times 10^{11}$, $\beta^*=0.30$ m, $\theta_c=348$ μ rad
- (5) “phase I w crab”, $N_b=1.6 \times 10^{11}$, $\beta^*=0.30$ m ($\theta_c=348$ μ rad)
- (6) “phase II+”, $N_b=2.3 \times 10^{11}$, $\beta^*=0.14$ m, $\theta_c=509$ μ rad
- (7) “phase II w (&w/o) crab”, $N_b=1.6 \times 10^{11}$, $\beta^*=0.14$ m, 509 μ rad
- (8) “LPA”, 50 ns, $N_b=4.2 \times 10^{11}$, $\beta^*=0.25$ m, $\theta_c=381$ μ rad
- (9) “LPA”, 25 ns, $N_b=2.6 \times 10^{11}$, $\beta^*=0.50$ m, $\theta_c=339$ μ rad

parameter	symbol	nom.	nom.*	ult.	$\beta^*=30$ cm, HI	$\beta^*=30$,cm , CC	$\beta^*=14$, cm HI	$\beta^*=14$ cm, CC	LPA – 25	LPA – 50
transverse emittance	ε [μm]	3.75	3.75	3.75	3.75	3.75	3.75	3.75	3.75	3.75
protons per bunch	N_b [10^{11}]	1.15	1.7	1.7	2.3	1.6	2.3	1.6	2.6	4.2
bunch spacing	Δt [ns]	25	50	25	25	25	25	25	25	50
beam current	I [A]	0.58	0.43	0.86	1.16	0.81	1.16	0.81	1.32	1.06
longitudinal profile		Gauss	Gauss	Gauss	Gauss	Gauss	Gauss	Gauss	Flat	Flat
rms bunch length	σ_z [cm]	7.55	7.55	7.55	7.55	7.55	7.55	7.55	11.8	11.8
beta* at IP1&5	β^* [m]	0.55	0.55	0.5	0.30	0.30	0.14	0.14	0.50	0.25
full crossing angle	θ_c [μrad]	285	285	315	348	(348)	509	(509)	339	381
Piwinski parameter	$\phi = \theta_c \sigma_z / (2^* \sigma_x^*)$	0.65	0.65	0.75	1.1	0.0	2.3	0.0	2.0	2.0
tune shift	ΔQ_{tot}	0.009	0.0136	0.009	0.01	0.01	0.006	0.01	0.01	0.01
peak luminosity	L [$10^{34} \text{ cm}^{-2}\text{s}^{-1}$]	1	1.1	2.3	5.9	4.0	7.5	7.9	4.0	7.4
peak events per #ing		19	40	44	111	76	142	150	75	280
initial lumi lifetime	τ_L [h]	23	16	15	7.7	7.8	6.0	4.0	12.4	5.3
effective luminosity ($T_{\text{turnaround}}=10 \text{ h}$)	L_{eff} [$10^{34} \text{ cm}^{-2}\text{s}^{-1}$]	0.45	0.43	0.90	1.8	1.2	2.0	1.7	1.5	1.9
	$T_{\text{run,opt}}$ [h]	21.5	17.7	17.2	12.4	12.5	11.0	8.9	16.0	10.5
effective luminosity ($T_{\text{turnaround}}=2 \text{ h}$)	L_{eff} [$10^{34} \text{ cm}^{-2}\text{s}^{-1}$]	0.67	0.68	1.41	3.2	2.2	3.8	3.5	2.4	3.6
	$T_{\text{run,opt}}$ [h]	9.6	7.9	7.7	5.5	5.6	4.9	4.0	7.2	4.7
e-c heat SEY=1.3	P [W/m]	0.4	0.1	0.6	1.3	0.7	1.3	0.7	1.4	0.8
SR heat 4.6-20 K	P_{SR} [W/m]	0.17	0.13	0.25	0.34	0.24	0.34	0.24	0.38	0.31
image current heat	P_{IC} [W/m]	0.15	0.17	0.33	0.60	0.29	0.60	0.29	0.39	0.51
gas-s. 100 h τ_b	P_{gas} [W/m]	0.04	0.03	0.06	0.08	0.05	0.08	0.05	0.09	0.07
luminous region	σ_1 [cm]	4.5	4.5	4.3	3.7	5.3	2.2	5.3	5.2	3.8
annual luminosity	L_{int} [fb $^{-1}$]	57	56	116	245	169	286	253	198	274

parameter highlights

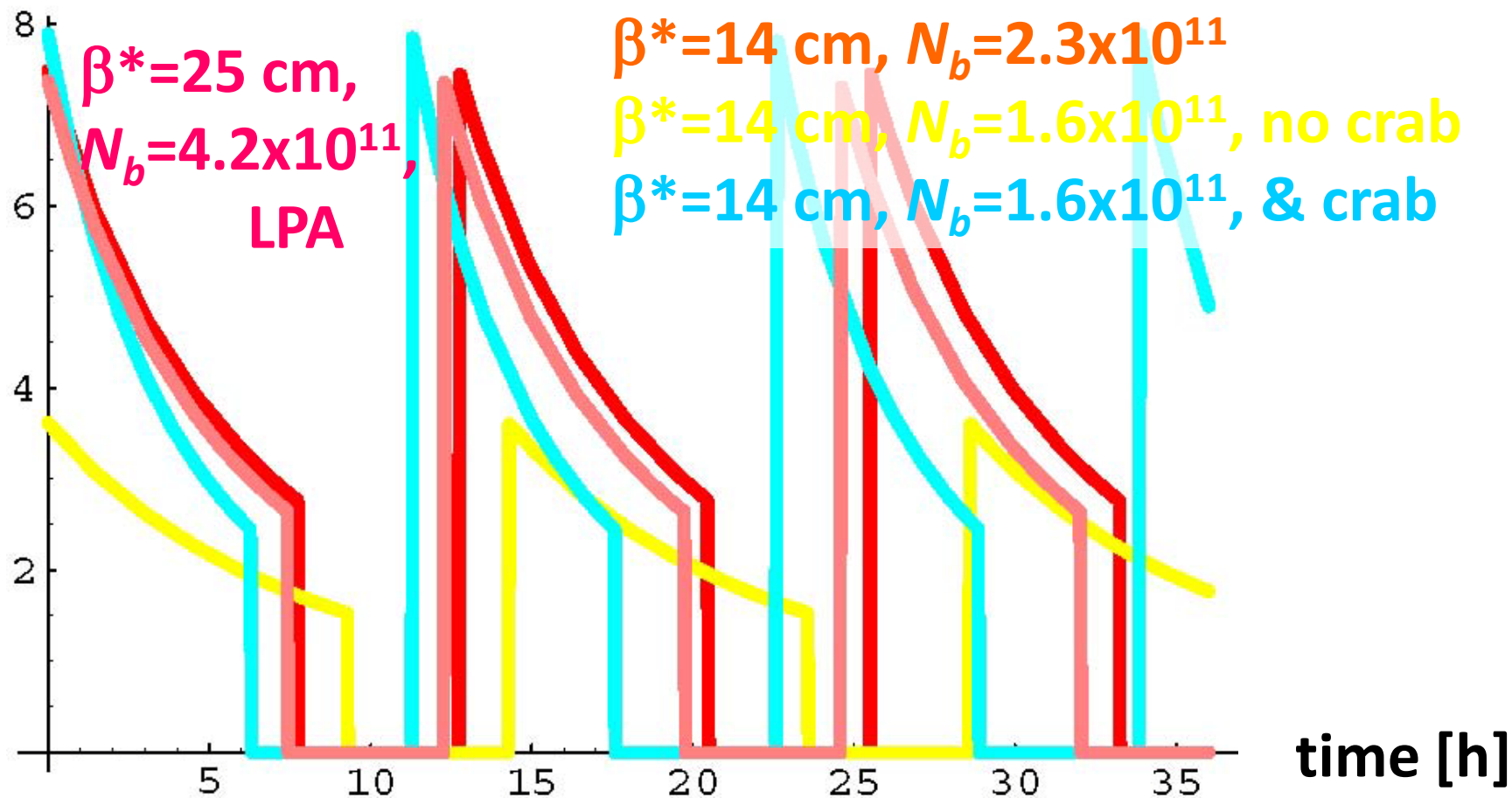
parameter	symbol	nom.	nom.* (50 ns).	ult.	$\beta^*=30$ (crab)	$\beta^*=14$	$\beta^*=14$ (crab)	LPA(50 ns, flat)
ppb	$N_b [10^{11}]$	1.15	1.7	1.7	1.6	2.3	1.6	4.2
beta* at IP1&5	$\beta^* [m]$	0.55	0.55	0.5	0.30	0.14	0.14	0.25
Piwinski angle		0.65	0.65	0.75	0.0	2.3	0.0	2.0
tune shift	ΔQ_{tot}	0.009	0.0136	0.009	0.01	0.006	0.01	0.01
peak luminosity	$L [10^{34}$ $cm^{-2}s^{-1}]$	1	1.1	2.3	4.0	7.5	7.9	7.4
peak evt's / #ing		19	40	44	76	142	150	280
lumi lifetime	$\tau_L [h]$	23	16	15	7.8	6.0	4.0	5.3
average ($T_{turnaround}=5$ h)	$L_{eff} [10^{34}$ $cm^{-2}s^{-1}]$	0.55	0.54	1.12	1.6	2.8	2.4	2.6
	$T_{run,opt} [h]$	15.2	7.9	12.2	8.8	7.7	6.3	7.5
annual lum. (200 days, 60% availability)	$L_{int} [fb^{-1}]$	57	56	116	168	286	253	274

luminosity evolution - examples



luminosity evolution – selected cases

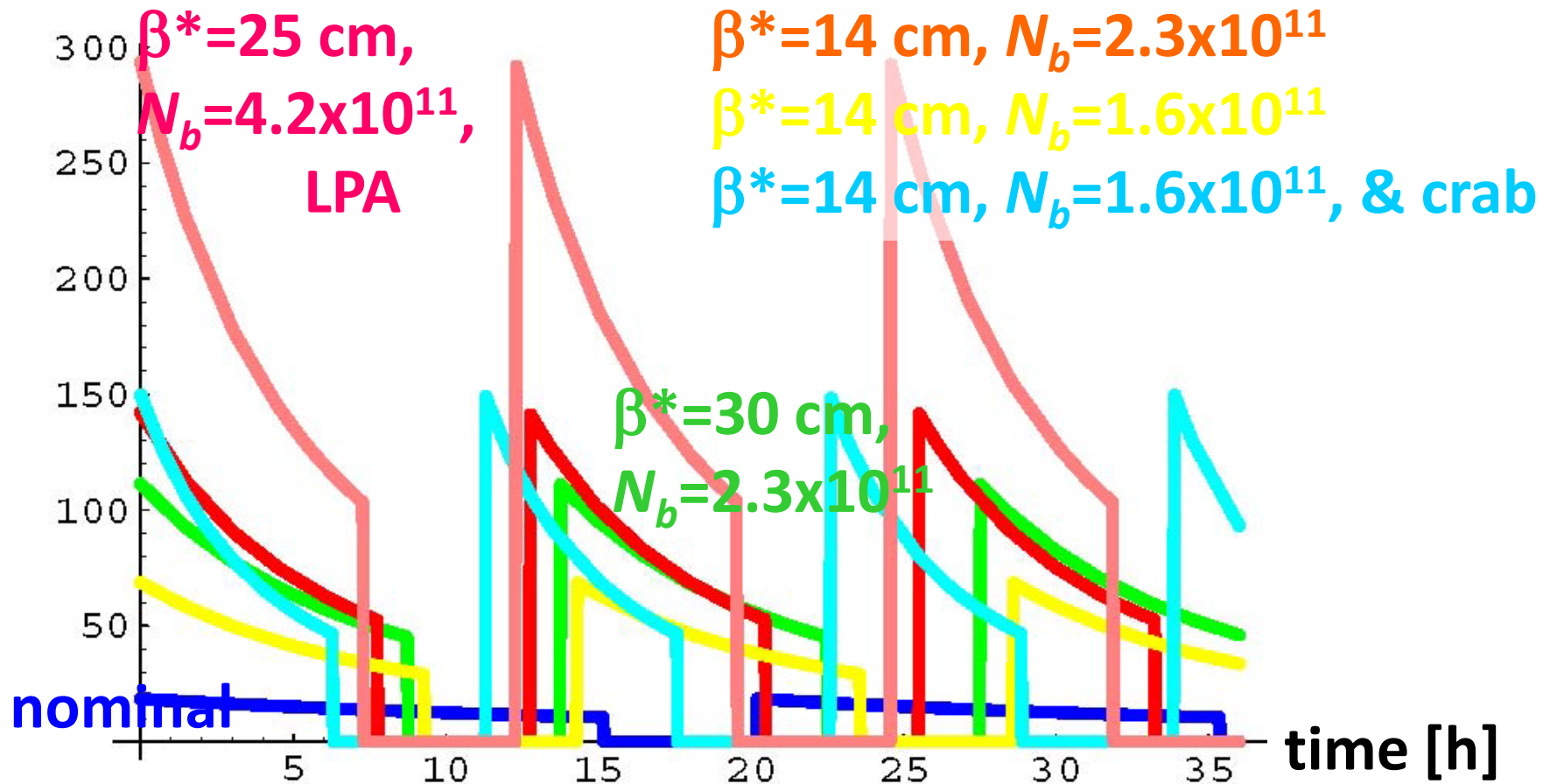
luminosity [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]



$\beta^*=14 \text{ cm} \ \& \ N_b=2.3 \times 10^{11}$ has very similar performance to $\beta^*=14 \text{ cm}, \ \& \ N_b \sim 1.6 \times 10^{11}$ and crab, and to $\beta^*=25 \text{ cm} \ \& \ N_b=4.2 \times 10^{11}$ & 50 ns spacing

events/crossing evolution

#events/crossing



all scenarios give peak #events/#ing $\sim 100-150$,
except for LPA ~ 300

luminosity leveling

changing θ_c , β^* or σ_z during the store in order to
→ **reduce event pile up & IR peak power deposition**
→ **maximize integrated luminosity**

leveling with crossing angle has two advantages:

increased average luminosity, operational simplicity

natural option for early separation or crab cavities,

leveling may first be tested in LHC heavy-ion collisions

two leveling strategies:

(1) constant luminosity

(2) constant beam-beam tune shift

optimum run time & av. luminosity

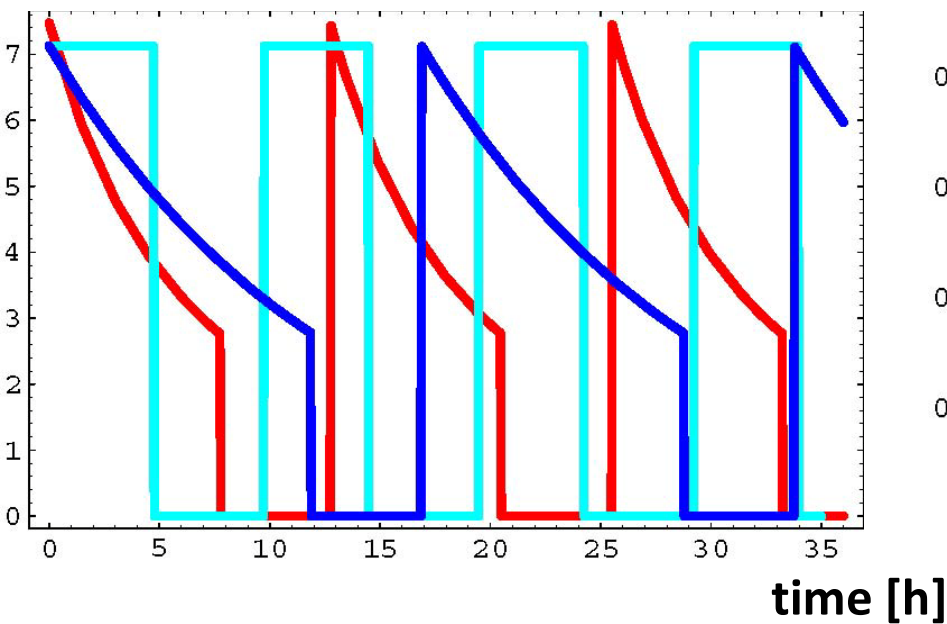
	w/o leveling	$L=\text{const}$	$\Delta Q_{\text{bb}}=\text{const}$
luminosity evolution	$L(t) = \frac{\hat{L}}{(1+t/\tau_{\text{eff}})^2}$	$L = L_0 \approx \text{const}$	$L(t) = \hat{L} \exp(-t/\tau_{\text{eff}})$
beam current evolution	$N(t) = \frac{N_0}{(1+t/\tau_{\text{eff}})}$	$N = N_0 - \frac{N_0}{\tau_{\text{eff}}} t$	$N(t) = N(0) \exp(-t/\tau_{\text{eff}})$
optimum run time	$T_{\text{run}} = \sqrt{\tau_{\text{eff}} T_{\text{ta}}}$	$T_{\text{run}} = \frac{\Delta N_{\text{max}} \tau_{\text{eff}}}{N_0}$	$T_{\text{run}} = \tau_{\text{eff}}$ $\min \left[\ln \left(\sqrt{1 + \phi_{\text{piw}}(0)^2} \right), \ln \left((T_{\text{ta}} + T_{\text{run}} + \tau_{\text{eff}}) / \tau_{\text{eff}} \right) \right]$
average luminosity	$L_{\text{ave}} = \hat{L} \frac{\tau_{\text{eff}}}{(\tau_{\text{eff}}^{1/2} + T_{\text{ta}}^{1/2})^2}$	$L_{\text{ave}} = \frac{L_0}{1 + \frac{L_0 \sigma_{\text{tot}} n_{\text{IP}} T_{\text{ta}}}{\Delta N_{\text{max}} n_b}}$	$L_{\text{ave}} = \frac{\tau_{\text{eff}}}{T_{\text{ta}} + T_{\text{run}}} \left(1 - e^{-T_{\text{run}}/\tau_{\text{eff}}} \right)$

leveling 2 \rightarrow exponential L decay, w decay time τ_{eff} (not $\tau_{\text{eff}}/2$)

leveling – example evolution

$$\beta^* = 14 \text{ cm}, N_b = 2.3 \times 10^{11}, T_{ta} = 5 \text{ h}$$

luminosity [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]

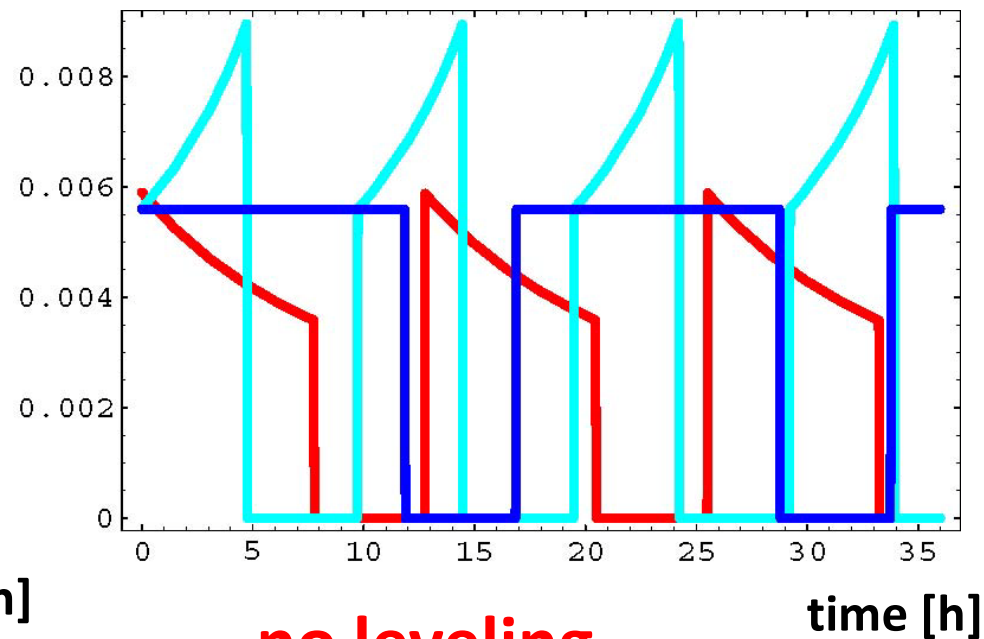


no leveling

$\Delta Q = \text{const}$

$L = \text{const}$

$|\Delta Q|$



no leveling

$\Delta Q = \text{const}$

$L = \text{const}$

leveling – example numbers

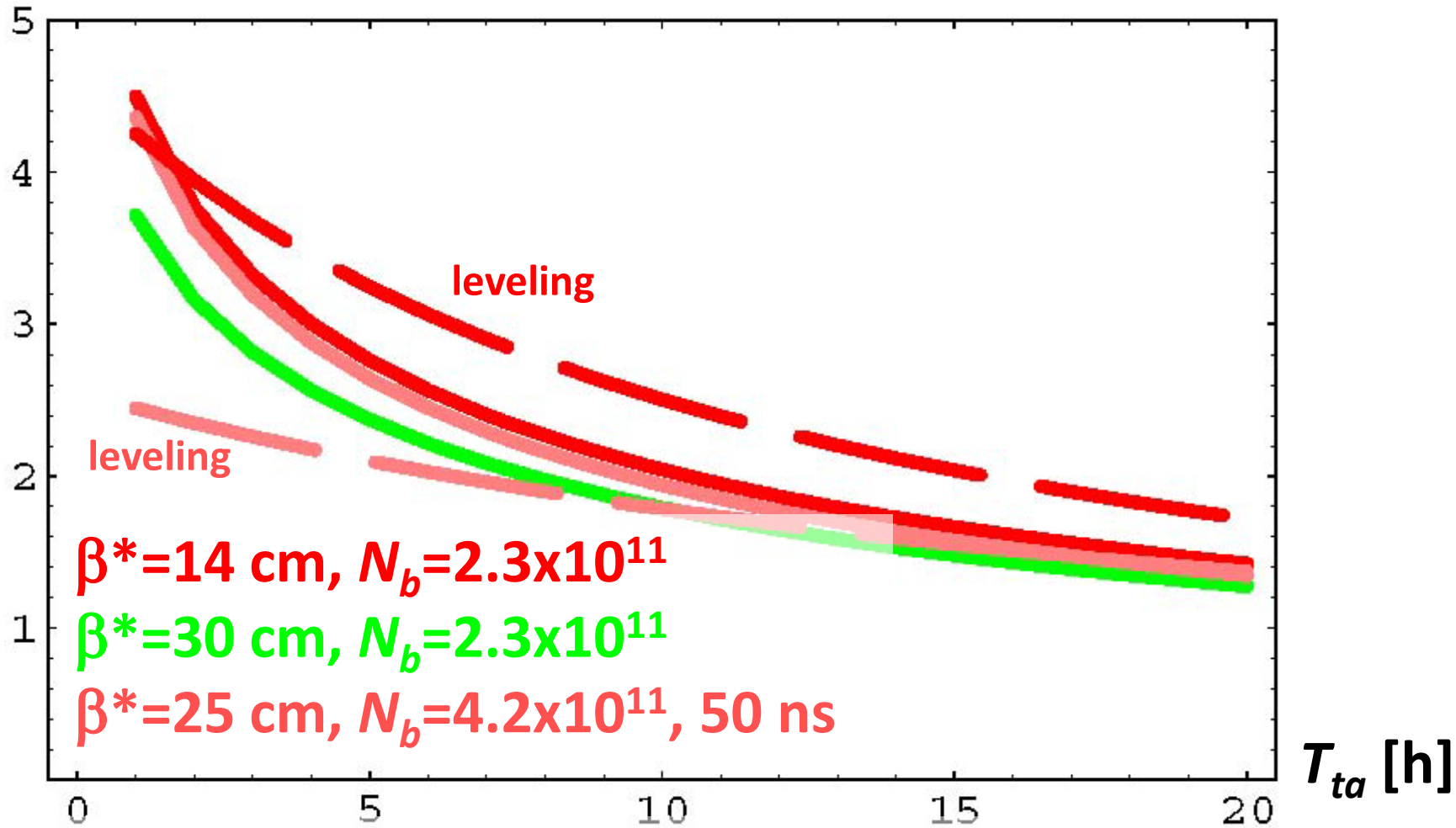
	$\beta^*=14$ cm, 25 ns spacing, $T_{ta}=5$ h			
	no leveling	$L=\text{const}$	$\Delta Q_{bb}=\text{const}$	
$N_b(0)$ [10^{11}]	2.3	2.3	2.3	2.3
$L(0)$ [$10^{34}\text{cm}^{-2}\text{s}^{-1}$]	7.5	7.1	12.3	7.1
$ \Delta Q_{bb}(0) $	0.0059	0.0056	0.01	0.0056
$ \Delta Q_{bb}(T_{run}) $	0.0036	0.0090	0.01	0.0056
$\theta_c(0)$ [μrad]	509	539	239	539
run time T_{run} [h]	7.74	4.74	2.72	11.9
$\langle L \rangle$ [$10^{34}\text{cm}^{-2}\text{s}^{-1}$]	2.8	3.5	3.6	3.2
events/#ing (0)	149	135	234	35

leveling – other example numbers

	$\beta^*=25$ cm, 50 ns spac., “LPA” $T_{ta}=5$ h		
	no leveling	$L=\text{const}$	$\Delta Q_{bb}=\text{const}$
$N_b(0)$ [10^{11}]	4.2	4.2	4.2
$L(0)$ [$10^{34}\text{cm}^{-2}\text{s}^{-1}$]	7.4	4.5	4.5
$ \Delta Q_{bb}(0) $	0.010	0.0056	0.0056
$ \Delta Q_{bb}(T_{run}) $	0.006	0.010	0.0056
$\theta_c(0)$ [μrad]	231	672	672
run time T_{run} [h]	7.45	6.0	23.2
$\langle L \rangle$ [$10^{34}\text{cm}^{-2}\text{s}^{-1}$]	2.6	2.5	2.4
events/#ing (0)	280	172	172

$\langle L \rangle$ vs. turnaround time

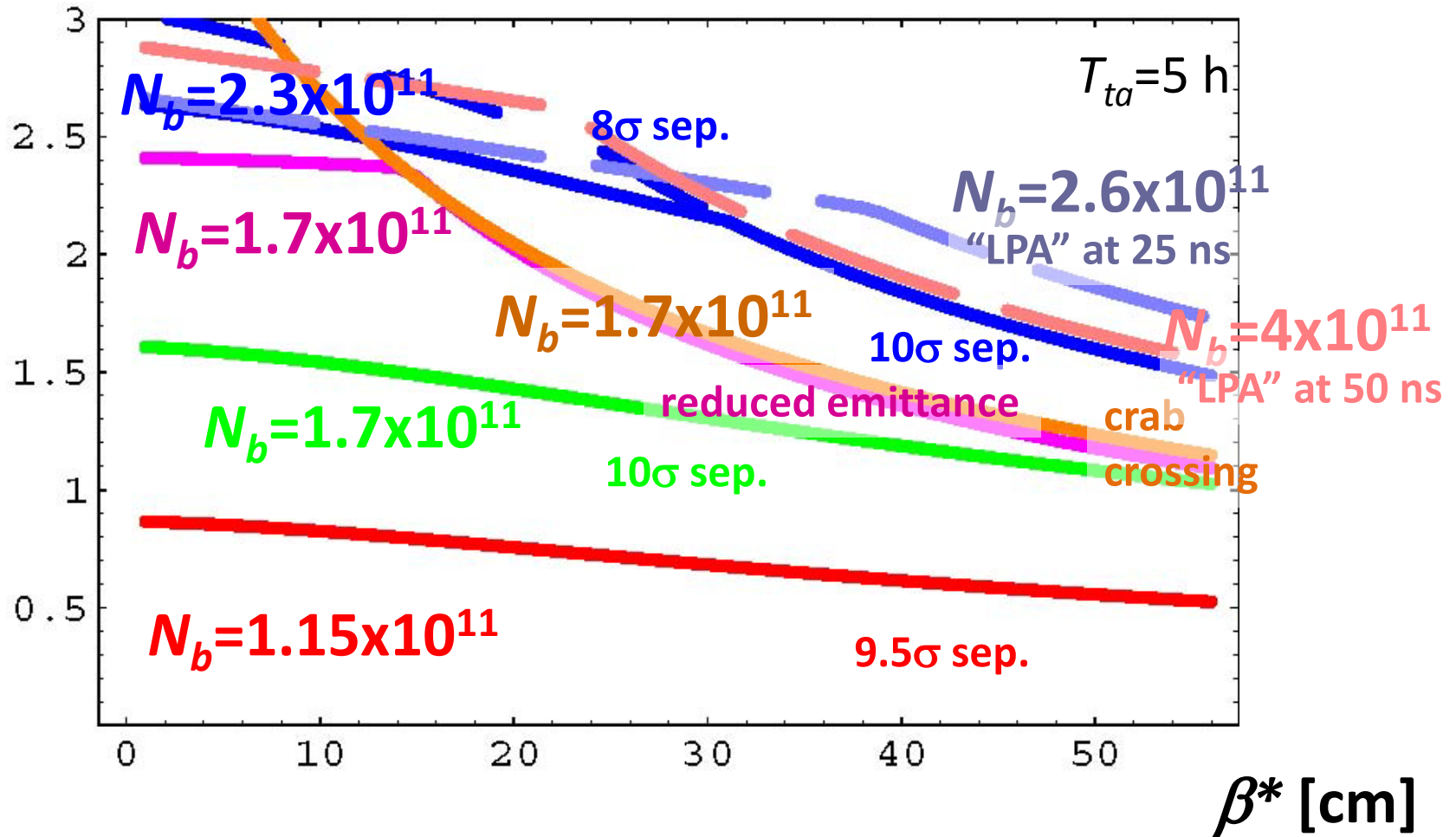
$\langle L \rangle$ [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]



reducing T_{ta} from 10 to 2 h increases $\langle L \rangle$ about 2x,
similar average luminosity for all 3 scenarios

$\langle L \rangle$ vs. β^* - the KEY PLOT

$\langle L \rangle$ [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]



beam intensity is much more important than β^* , reducing β^* only helps in the presence of crab cavities

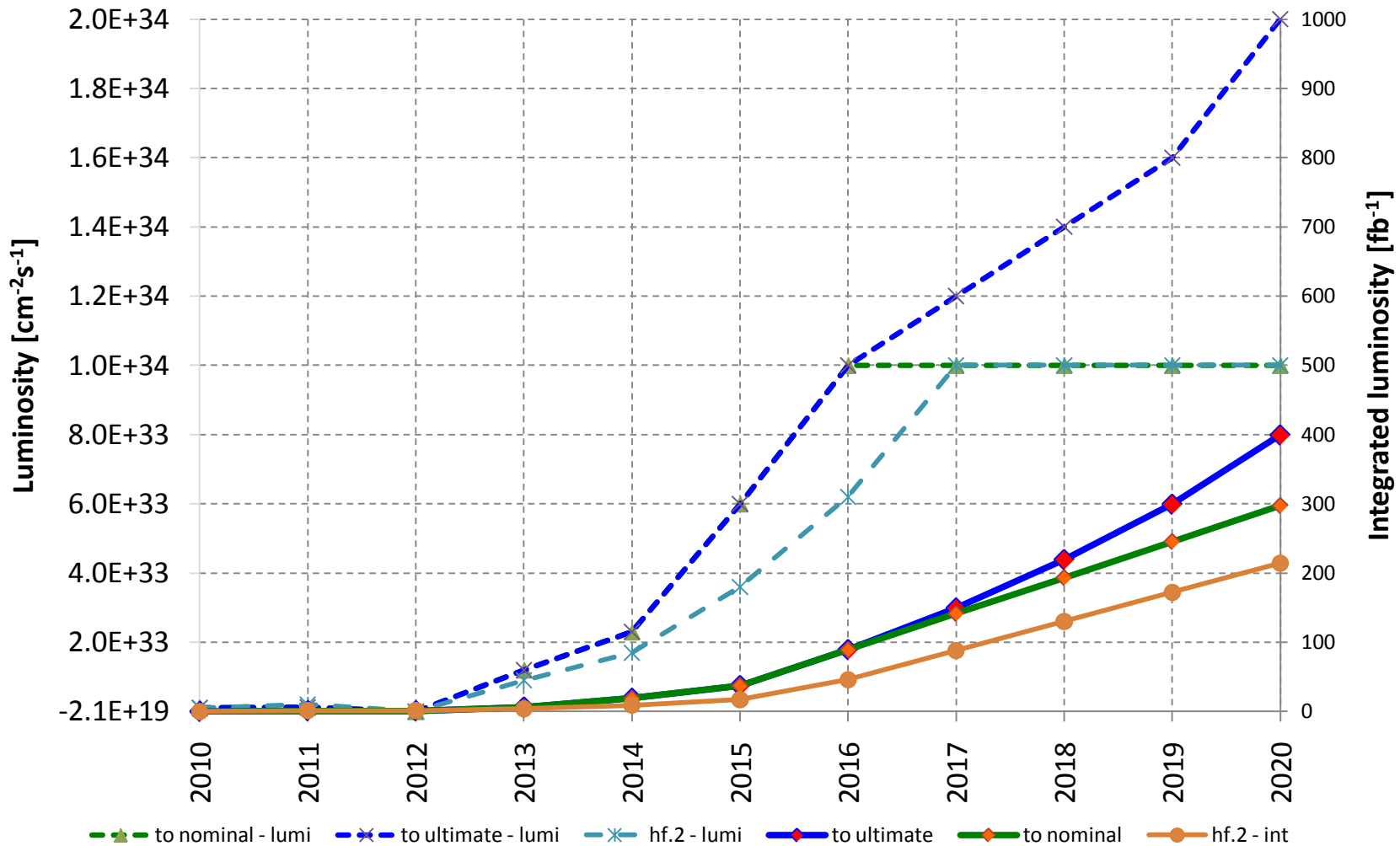
conclusions

- upgrade scenarios with 25 & 50 ns spacing
- maximum $N_b \sim 2.3 \times 10^{11}$ at 25 ns, $\sim 5.0 \times 10^{11}$ at 50 ns
- T_{ta} - 10 \rightarrow 2 h: 2x higher $\langle L \rangle$
- β^* : factor 2 reduction \rightarrow 10-20% higher $\langle L \rangle$
- N_b : factor 2 increase \rightarrow 3 times higher $\langle L \rangle$!
- crab crossing: 20-100% higher $\langle L \rangle$
- luminosity optimization assumes two IPs;
needs/policy for ALICE & LHCb?
- θ_c leveling can increase run time by factor 1.5-3,
& reduce pile up, at \sim constant $\langle L \rangle$
- annual luminosities of 150-300 fb⁻¹
- put emphasis on N_b (!!), T_{ta} (!) and crab crossing

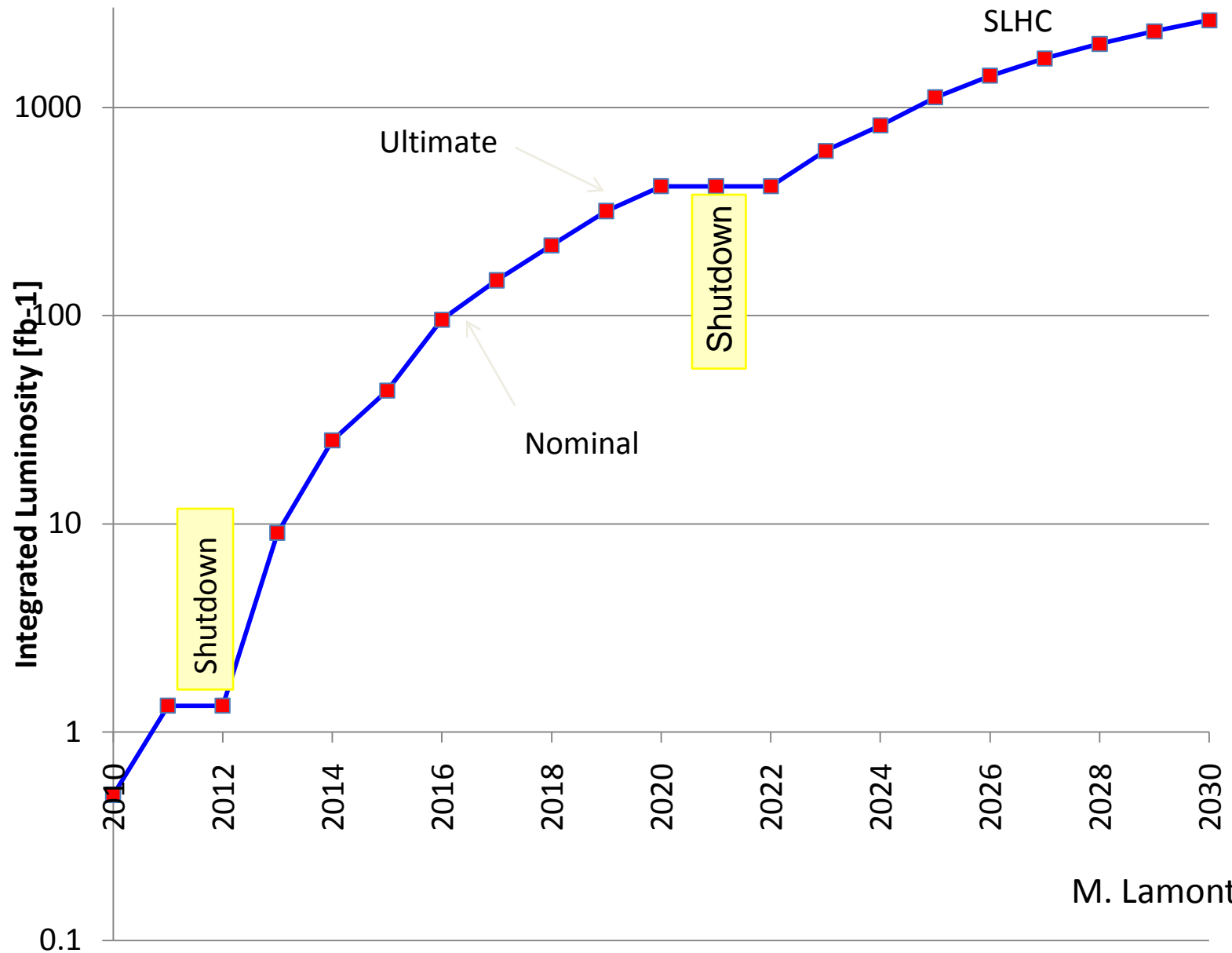
upgrade time lines

- **collimation** upgrade 2013-2015
- parallel R&D on Nb-Ti and Nb₃Sn **IR quadrupoles**
technology choice by 2014
- parallel R&D on compact and global **crab cavities**
choice by 2014
- **LINAC4** connection 2014-15
- **injector** upgrades & consolidation (PSB, PS, SPS)
completed by ~2015
- production & installation of **new IRs** ~2018-22
- production & installation of **crab cavities** ~2018-22

possible luminosity evolution → 2020



possible luminosity evolution → 2030



thank you for your attention!