



CMS Experiment at the LHC, CERN at 2.36 TeV

Data recorded: 2009-Dec-14 04:05:38.307318 GMT

Run: 124120

Event: 945133

Lumi section: 31

Orbit: 31924351

Crossing: 51

"Phase-2" Scenarios

Frank Zimmermann

LHCC Upgrade Review

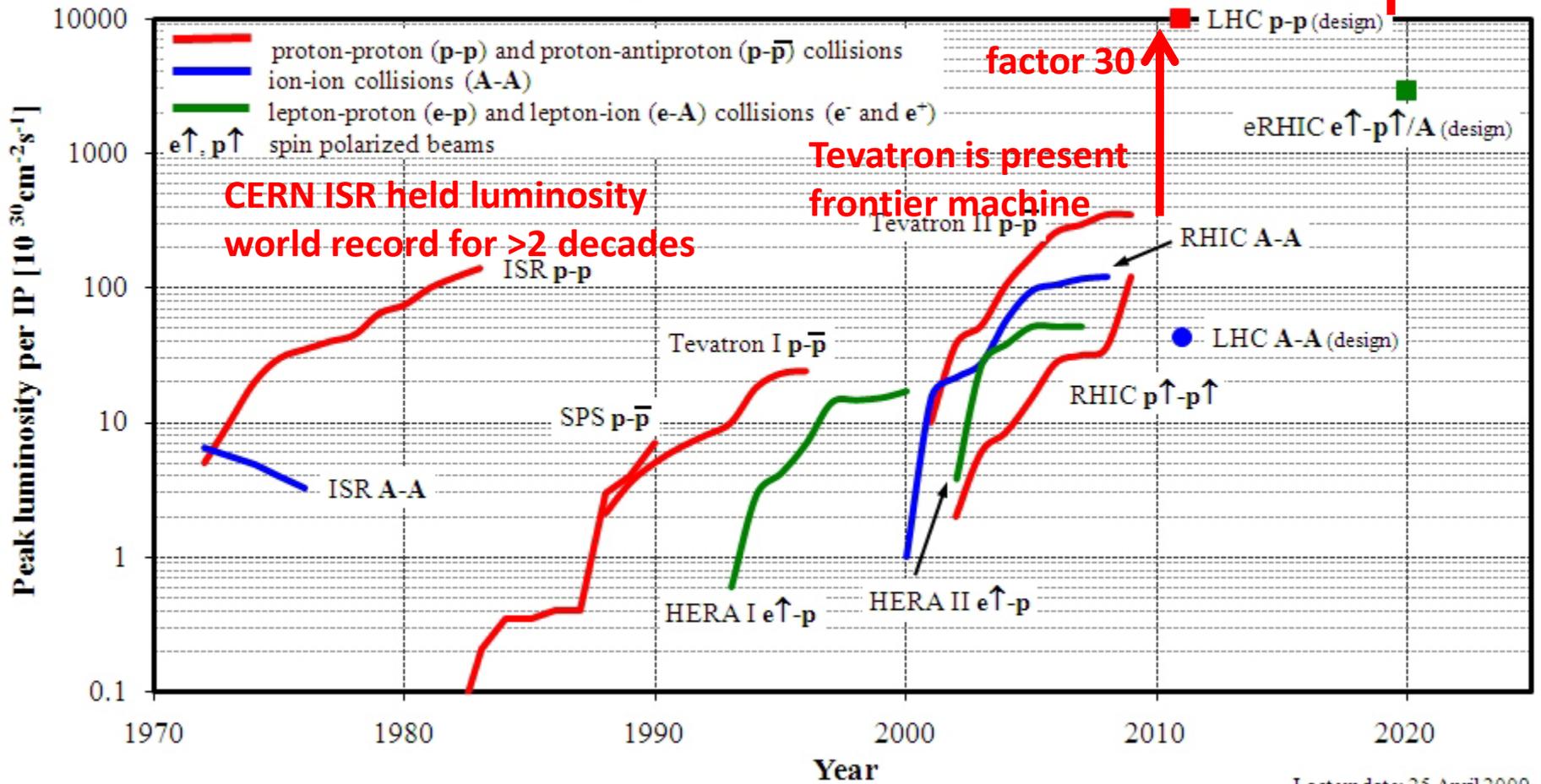
February 2010

input from 2001 LHC Upgrade Feasibility Study and from numerous
CARE-HHH and EuCARD-AccNet workshops



special thanks to R. Assmann, R. Bailey, C. Bhat, O. Brüning, R. Calaga, H. Damerou, D. Denegri, O. Dominguez, U. Dorda, L. Evans, S. Fartoukh, R. Garoby, M. Giovannozzi, B. Goddard, N. Hessey, B. Jeanneret, E. Jensen, J.-P. Koutchouk, H. Maury Cuna, S. Myers, M. Nessi, K. Ohmi, R. Ostojic, Y. Papaphilippou, L. Rossi, F. Ruggiero, G. Rumolo, W. Scandale, D. Schulte, E. Shaposhnikova, G. Sterbini, K. Takayama, L. Tavian, T. Taylor, E. Todesco, R. Tomas and E. Tsesmelis

Luminosity evolution of hadron colliders



Courtesy W. Fischer

disclaimer

LHC upgrade plans & schedule under review at:

- LHC Machine Committee (weekly)
- special “brainstorming” meetings
- directorate retreat mid-November
- Chamonix 2010 workshop (Jan. ‘10)
- CERN MAC (1st mtg. 26 October)
- LHC “lumi up” task force (next week)

previous assumptions & schedules are likely to change significantly

plans, scenarios & time scales being revised...

contents of this presentation

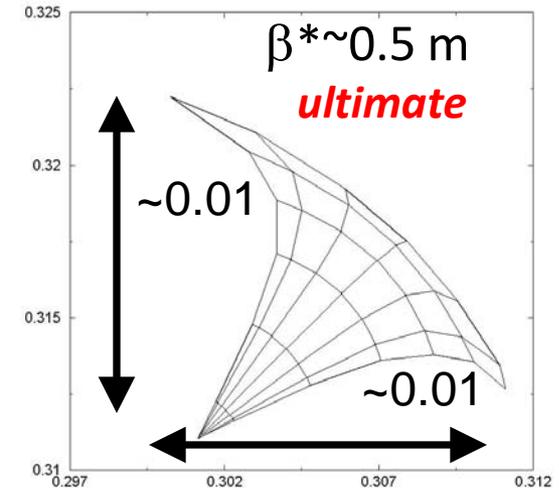
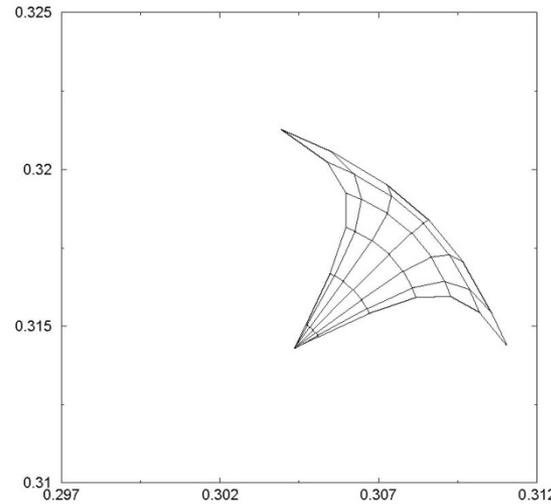
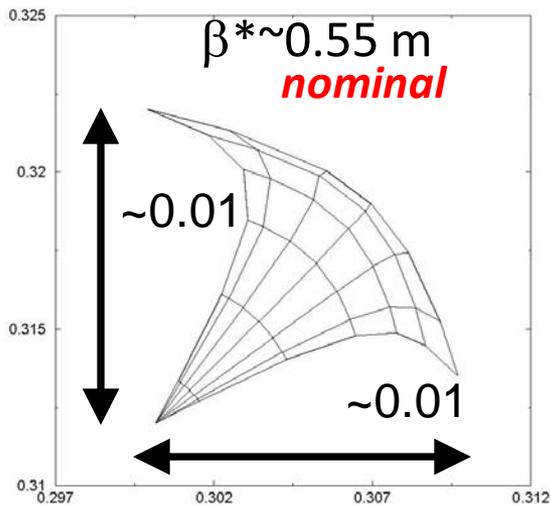
- 1) parameters
- 2) the original plan; LHCb & ALICE?
- 3) few words about phase-I
- 4) constraints & collision schemes
- 5) recent progress (CC, LPA, e-cloud)
- 6) example scenarios
- 7) luminosity leveling
- 8) turnaround time, β^* , intensity
- 9) conclusions & questions

parameters

- β^* - IP beta function
- β_x^*/β_y^* - ratio of IP beta functions
- θ_c - (full) crossing angle
- ε_N - normalized transverse emittance
- N_b - bunch intensity
- n_b - number of bunches ($\rightarrow s_b$ - bunch spacing)
- longitudinal bunch profile (“flat” vs “Gaussian”)
- number of collision points (IP’s)
- T_{ta} - turn-around time

#IP's : the original plan – “phase 0”

J.Gareyte, F. Ruggiero *et al*, e.g. LHC'99 workshop, LHC Project Report 626



nominal tune footprint
up to 6σ with **4 IPs & nom.**
intensity $N_b = 1.15 \times 10^{11}$

$$L = 10^{34} \text{ cm}^{-2}\text{s}^{-1}$$

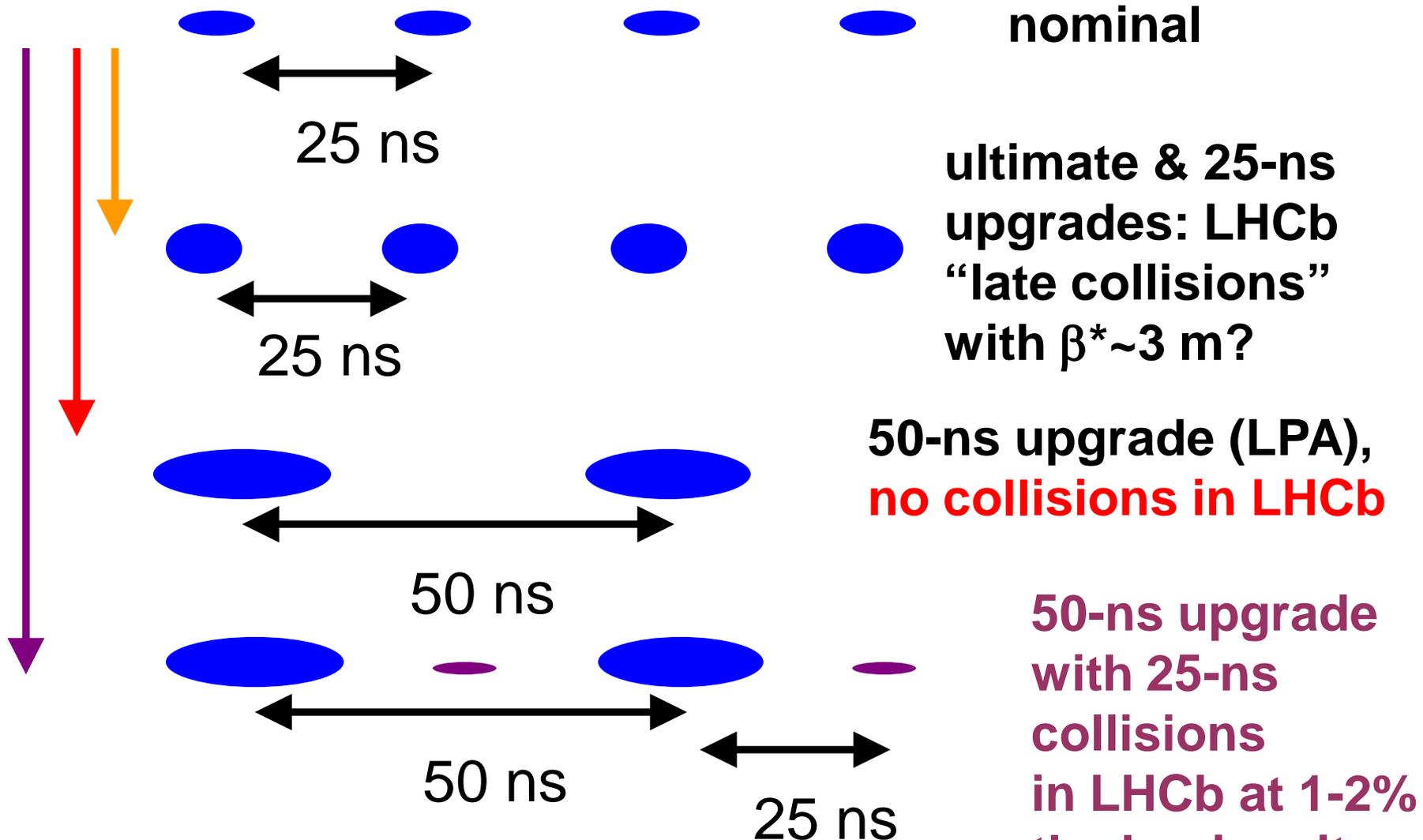
tune footprint up to 6σ
with nominal intensity
and **2 IPs**

tune footprint up to 6σ
with **2 IPs at ultimate**
intensity $N_b = 1.7 \times 10^{11}$

$$L = 2.3 \times 10^{34} \text{ cm}^{-2}\text{s}^{-1}$$

“going from 4 to 2 IPs ATLAS & CMS luminosity can be increased
by factor 2.3 - further, increasing crossing angle to $340 \mu\text{rad}$,
bunch length (x2), & bunch charge to $N_b = 2.6 \times 10^{11}$ would yield
 $L = 3.6 \times 10^{34} \text{ cm}^{-2}\text{s}^{-1}$ [$\beta^* = 0.5$ m]”

what about LHCb? – bunch patterns



nominal

ultimate & 25-ns
upgrades: LHCb
“late collisions”
with $\beta^* \sim 3$ m?

50-ns upgrade (LPA),
no collisions in LHCb

50-ns upgrade
with 25-ns
collisions
in LHCb at 1-2%
the luminosity

LHC-IR “phase-I”: merits & concerns

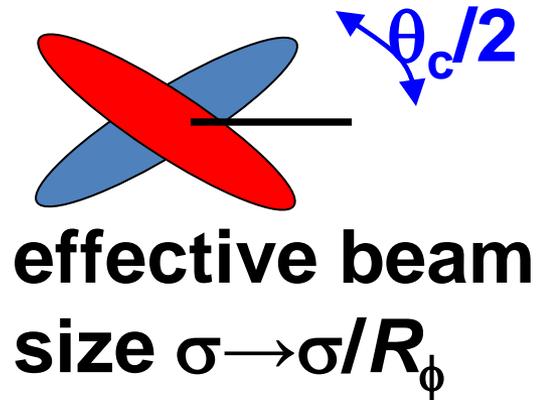
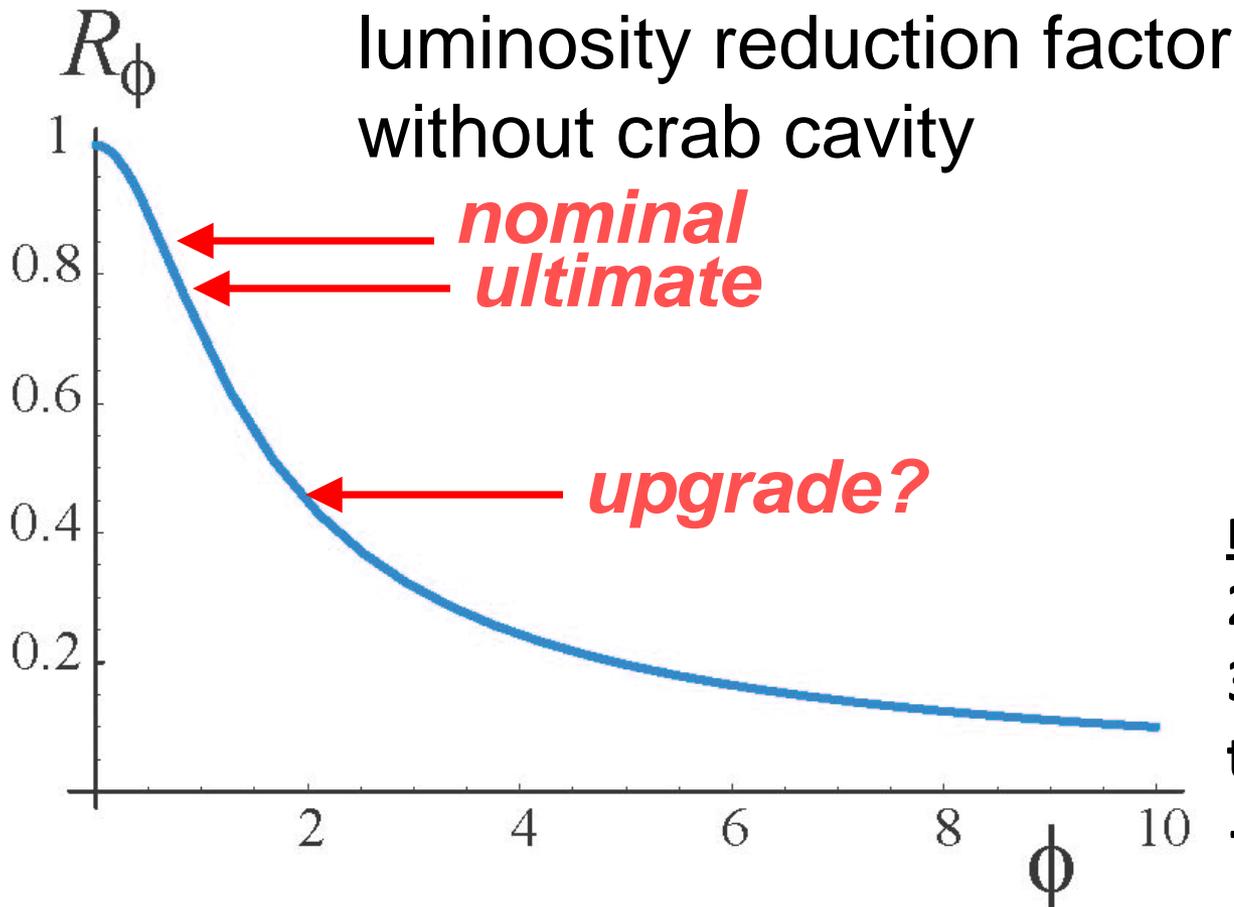
- + β^* reduction by up to a factor of 2
- + larger aperture in triplet
- potential loss in optics flexibility
- higher chromaticity & chromatic aberrations
- more parasitic long-range beam-beam collisions
- about 1 year downtime

upgrade constraints

- **total beam-beam tune shift ≤ 0.01**
 - SPS p-pbar experience
- long-range beam-beam \rightarrow **crossing angle $\geq 9\sigma$**
- **arc cooling capacity**
 - global & local limitations, cooling shares with IR
 - heat load from SR, image currents, & e-cloud
- IR layout & optics $\rightarrow \beta^*$
- **event pile up** in the detectors (≤ 300 , ≤ 150 ?)
- **luminosity lifetime** (≥ 5 h?)

constraint - crossing angle

$$R_\phi = \frac{1}{\sqrt{1 + \phi^2}}; \quad \phi \equiv \frac{\theta_c \sigma_z}{2\sigma_x^*} \quad \text{“Piwinski angle”}$$



range - $f(\text{triplet}, \beta^*)$:
285 μrad (nominal)
315 μrad (ultimate)
till $\sim 410 \mu\text{rad}$ “phase I”
 $\rightarrow 500 \mu\text{rad}$ “phase II”?

b-b tune shift, ϕ & luminosity

$$\Delta Q_{bb} = \frac{N_b}{\gamma \varepsilon} \frac{r_p}{2\pi} \frac{1}{\sqrt{1 + \phi_{piw}^2}} \frac{1}{F_{profile}}$$

total b-b tune shift
for two IP's with
alternating crossing

$$L = \frac{1}{4\pi} f_{rev} n_b \gamma \frac{1}{\beta^* (\gamma \varepsilon)} N_b^2 \frac{1}{\sqrt{1 + \phi_{piw}^2}}$$

at the b-b limit, larger Piwinski angle &/or larger emittance increase luminosity!

$$= \frac{\pi}{r_p^2} f_{rev} n_b \gamma \frac{(\gamma \varepsilon)}{\beta^*} \Delta Q_{bb}^2 F_{profile}^2 \sqrt{1 + \phi_{piw}^2}$$

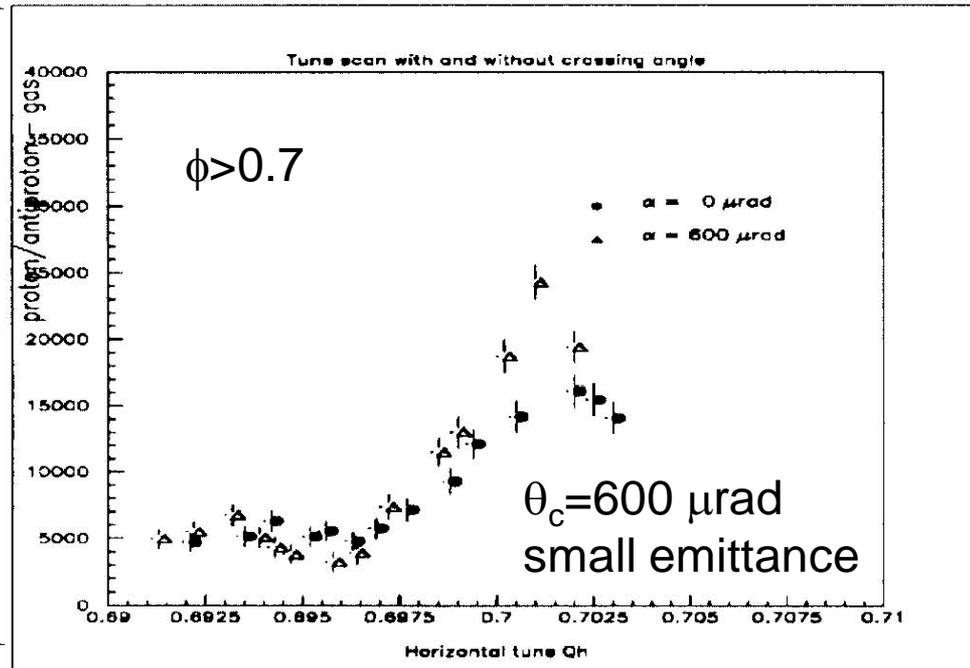
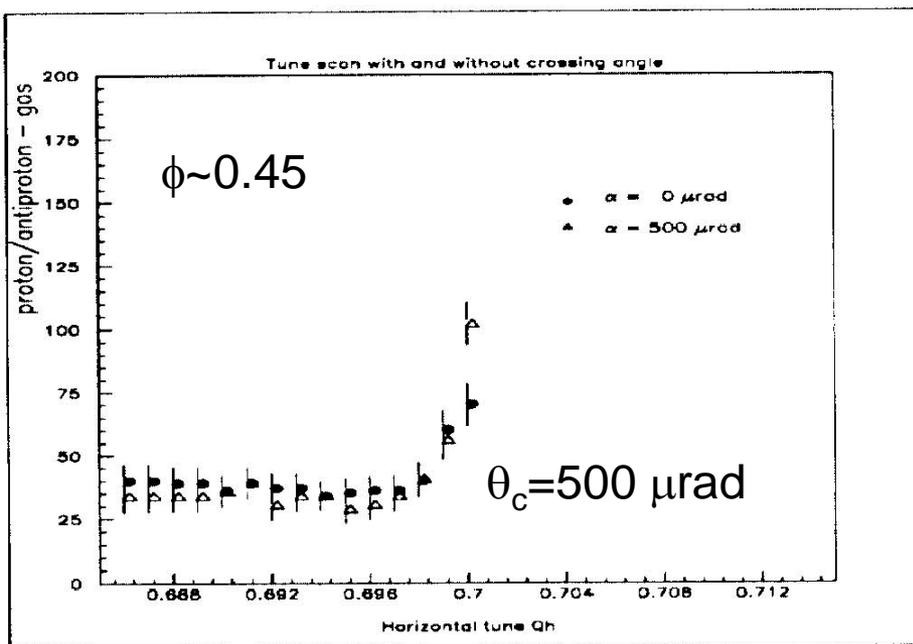
optimization strategies:

- 1) increase N_b with ε (e.g. controlled ε blow up at top energy)
- 2) increase N_b with $1/R_\phi$ & “flat” bunch $F_{profile} \sim 1.4$ (“LPA”)
- 3) vary ε as $1/R_\phi$ (“small emittance”)
- 4) set $1/R_\phi = 1$ at IP and minimize β^* (e.g. crab crossing)

beam-beam limit – θ_c dependence?

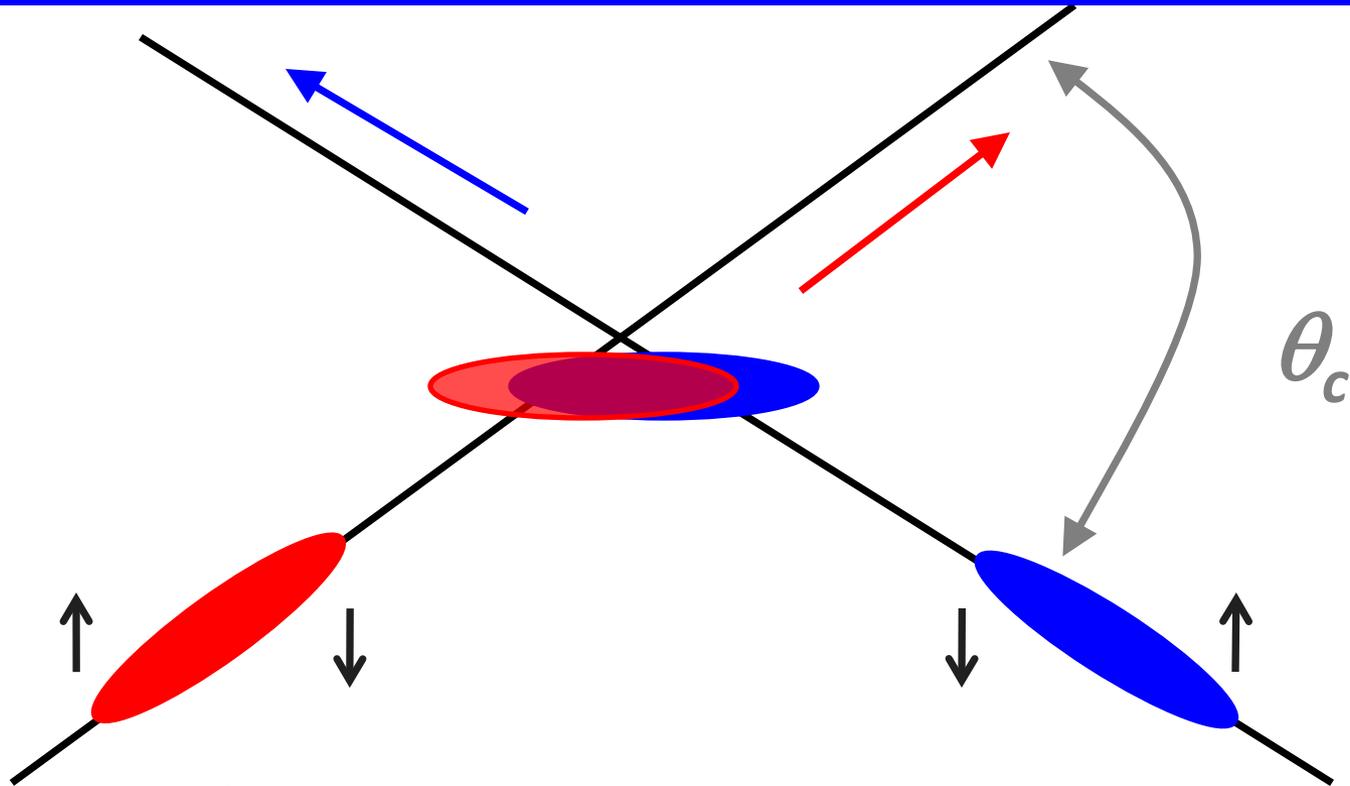
in lepton colliders crossing angle has reduced the beam-beam limit (DORIS-I, KEKB,...)

for hadrons, one historical experiment at the SPS
K. Cornelis, W. Herr, M. Meddahi, PAC91 San Francisco



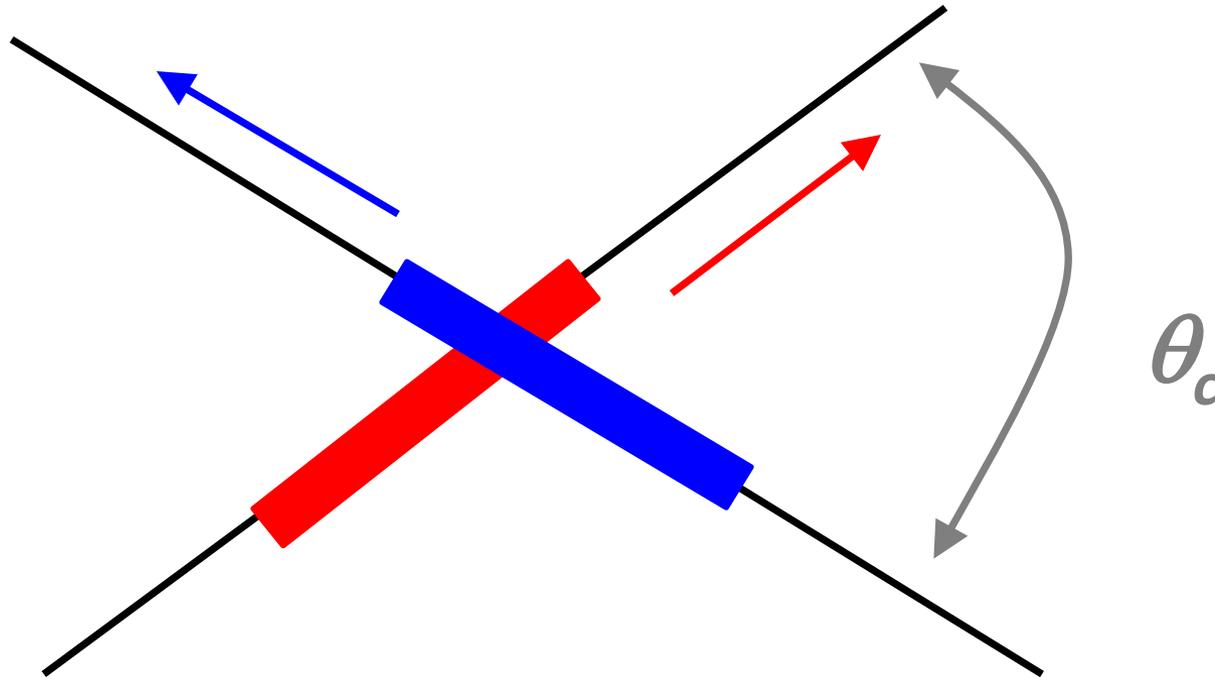
(almost) no additional beam-beam effect, but ϕ was much smaller than considered for SLHC

crab crossing



- RF crab cavity deflects head and tail in opposite direction so that collision is effectively “head on” for luminosity and tune shift
 - bunch centroids still cross at an angle (easy separation)
 - 1st proposed in 1988, in operation at KEKB since 2007
- advantages: higher geometric luminosity, easy leveling, potentially higher beam-beam tune shift**

large Piwinski angle – “LPA”



- 1) large Piwinski angle $\theta_c \sigma_z \gg 2 \sigma_x^*$
 - 2) longitudinally flat profile
- reduced tune shift, higher bunch charge
(& 50 ns spacing for e-cloud)**

recent progress on “phase-II” schemes

efforts focus on crab crossing & LPA scheme:

- ✓ crab cavities
- ✓ generation & stability of long flat bunches
- ✓ electron cloud simulations

LHC-CC09 workshop

**LHC Crab Cavity Workshop,
jointly organized by CERN,
EuCARD-ACCNET, US-LARP,
KEK, & Daresbury
Lab/Cockcroft Institute
CERN, 16-18 September 2009**



~50 participants, LHC Crab Cavity Advisory Board established



CERN statement (Steve Myers) on LHC crab cavities issued after AccNet LHC-CC09 workshop

Statements on Crab Cavities from CERN

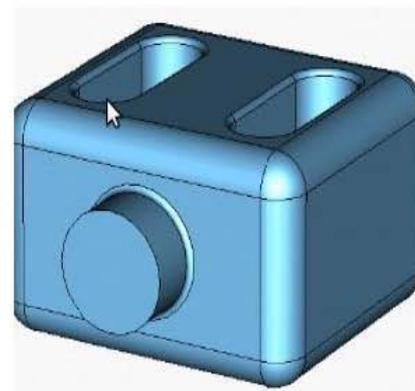
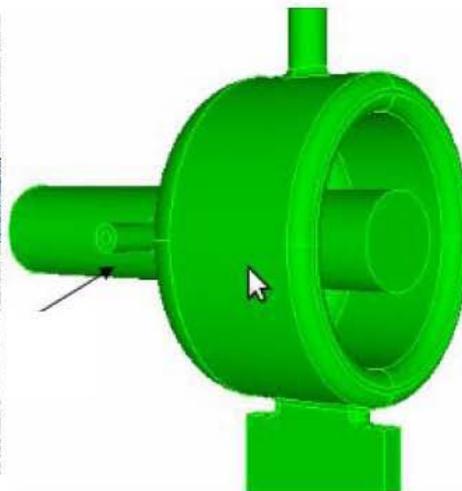
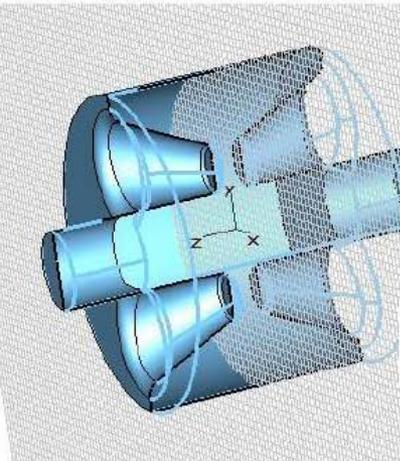
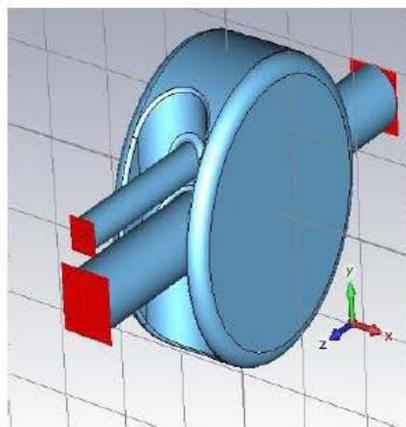
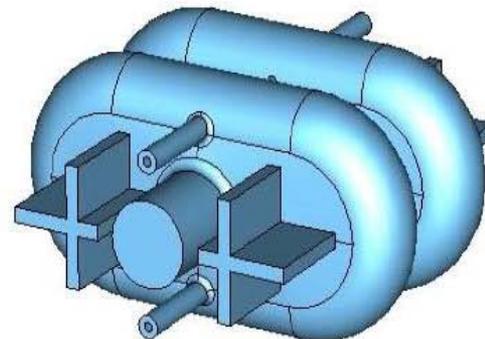
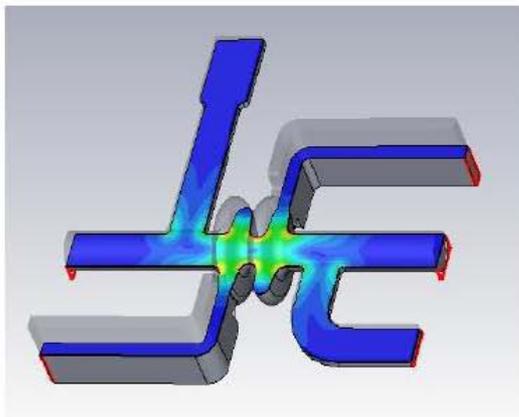
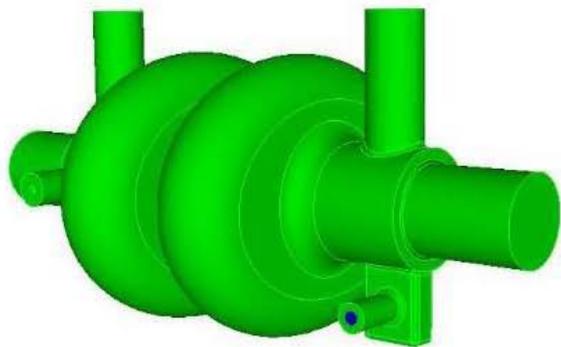
(Steve Myers, Director of Accelerators and Technology)

1. Following the success of KEKB, CERN must pursue the use of crab cavities for the LHC, since the potential luminosity increase is significant.
2. A final crab-cavity implementation for the LHC has not yet been settled. Both “local” and “global” crabbing schemes are still under consideration for the LHC upgrade phase II. Future R & D should focus on compact cavities which are suitable for both schemes.
3. One possible show-stopper has been highlighted: machine protection, which is critical for LHC. The effect of fast cavity changes needs to be looked at with high priority. Mitigation schemes such as raising the Q value of the cavity to $\sim 10^6$ (from $\sim 10^5$ at KEK) will be studied.
4. Another important issue is the impedance. Since the LHC revolution frequency changes during acceleration, the detuning of the cavity may be more difficult than was the case for KEKB, and other measures (like strong damping of the dipole mode) need to be examined.
5. High reliability of the crab cavities is essential; the trip rate should be low enough not to perturb LHC beam operation.
6. Validation cavity tests in the LHC itself are not deemed essential. It is considered plausible to install a new system in the LHC without having tested a prototype in the LHC beforehand. As in all new colliders, this has been done with many other components.
7. Demonstration experiments should focus on the differences between electrons and protons (e.g. effect of crab-cavity noise with beam-beam tune spread; impedance; beam loading) and on reliability & machine protection which are critical for the LHC.
8. A beam test with a KEKB crab cavity in another proton machine is considered useful, meaningful and sufficient (for deciding on a full crab-cavity implementation in LHC) if it addresses the differences between protons and electrons.
9. Possible modifications of LHC Interaction Region 4 during the 2013/14 shutdown should be studied to evaluate the feasibility of installing and testing crab-cavity prototypes, and of accommodating a possible global crab-cavity scheme.
10. The timing of the crab-cavity implementation should be matched to the short and long-term goals and to the overall CERN schedule, and be in phase with the experiment upgrades.
11. **The crab-cavity infrastructure should be included in all other LHC upgrades scenarios.**
12. **Crab cavities can increase the LHC luminosity without an accompanying increase in beam intensity, thereby avoiding negative side effects associated with high intensity and high stored beam energy. This opinion has been endorsed by the general-purpose high-luminosity experiments.**

CERN statements (excerpts)

1. **KEKB success ... CERN must pursue crab cavities for LHC**
2. ... Future R&D should focus on **compact cavities** ... suitable for both [local and global] schemes
7. **Demonstration experiments should focus on differences between electrons and protons** (e.g. effect of crab-cavity noise with beam-beam, impedance, beam loading) and on reliability & machine protection which are critical for LHC
8. **A beam test with KEBB crab cavity in another proton machine** ... useful, meaningful and sufficient ...
9. **Possible modifications of Interaction Region 4** during the 2013/14 shutdown
11. Crab cavity infrastructure ... **be included in all ... LHC upgrades**
12. **Crab cavities can increase luminosity w/o accompanying increase in beam intensity, thereby avoiding negative side effects**

CC designs presented at LHC-CC09



further crab cavity progress

30 October 2009:

launch of CERN working group on feasibility of KEKB crab cavity test in SPS

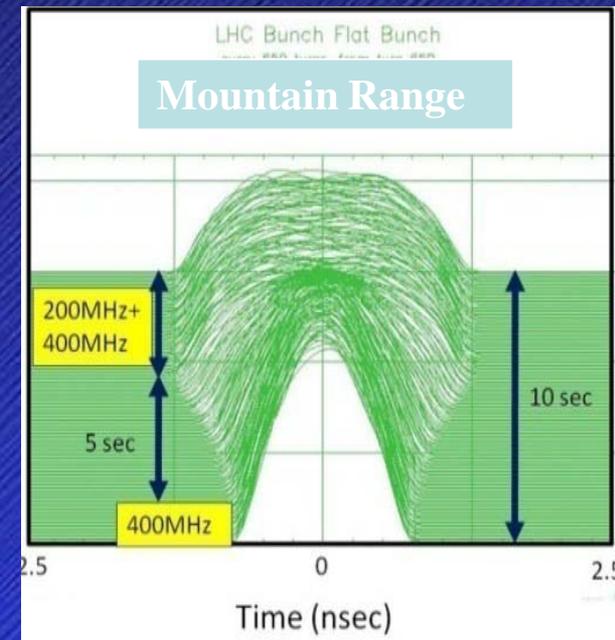
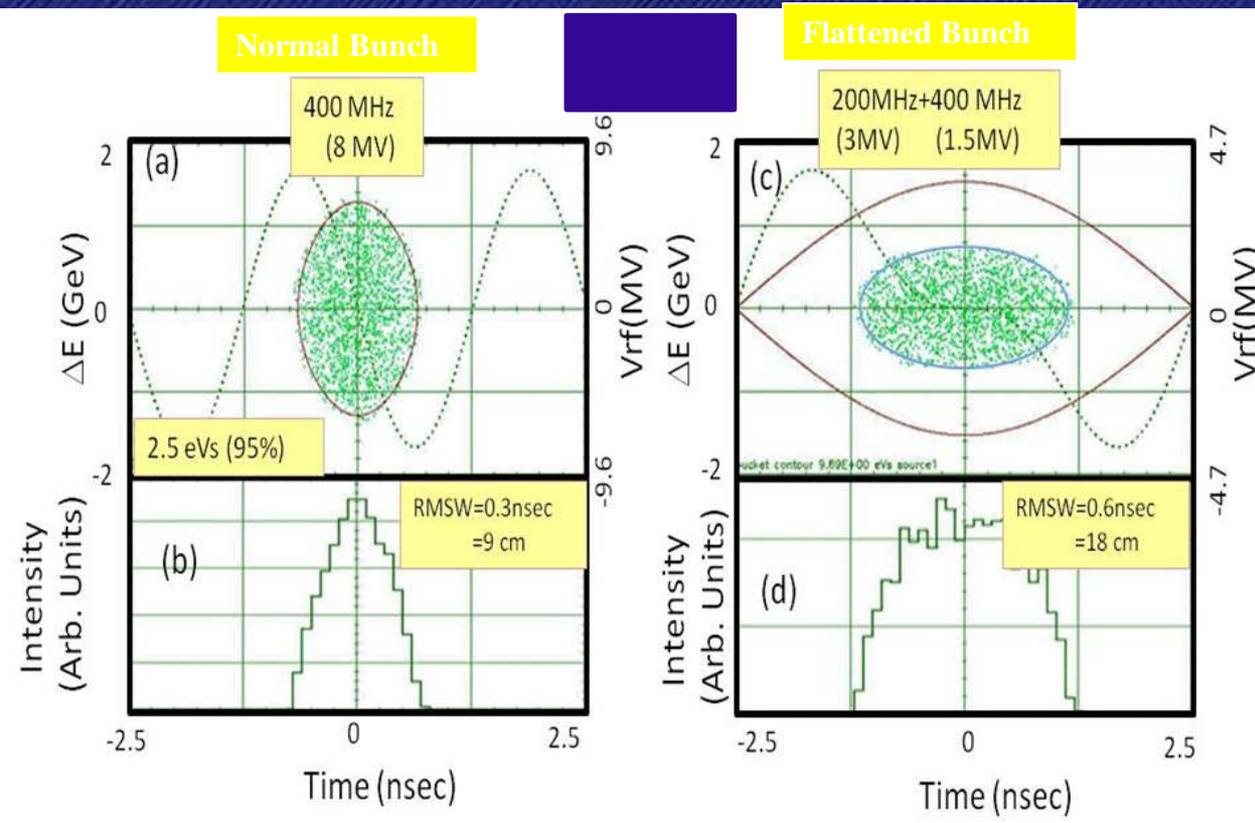
WG conclusions on 18 December 2009:

no real showstoppers; KEKB crab cavity could be used/tested at SPS in 2012; best location found (space & available cryogenics); SPS beam test including LHC collimators; effect of RF noise; trip rates; proposal of bypass (i.e. 2 movable beam pipes w Y)

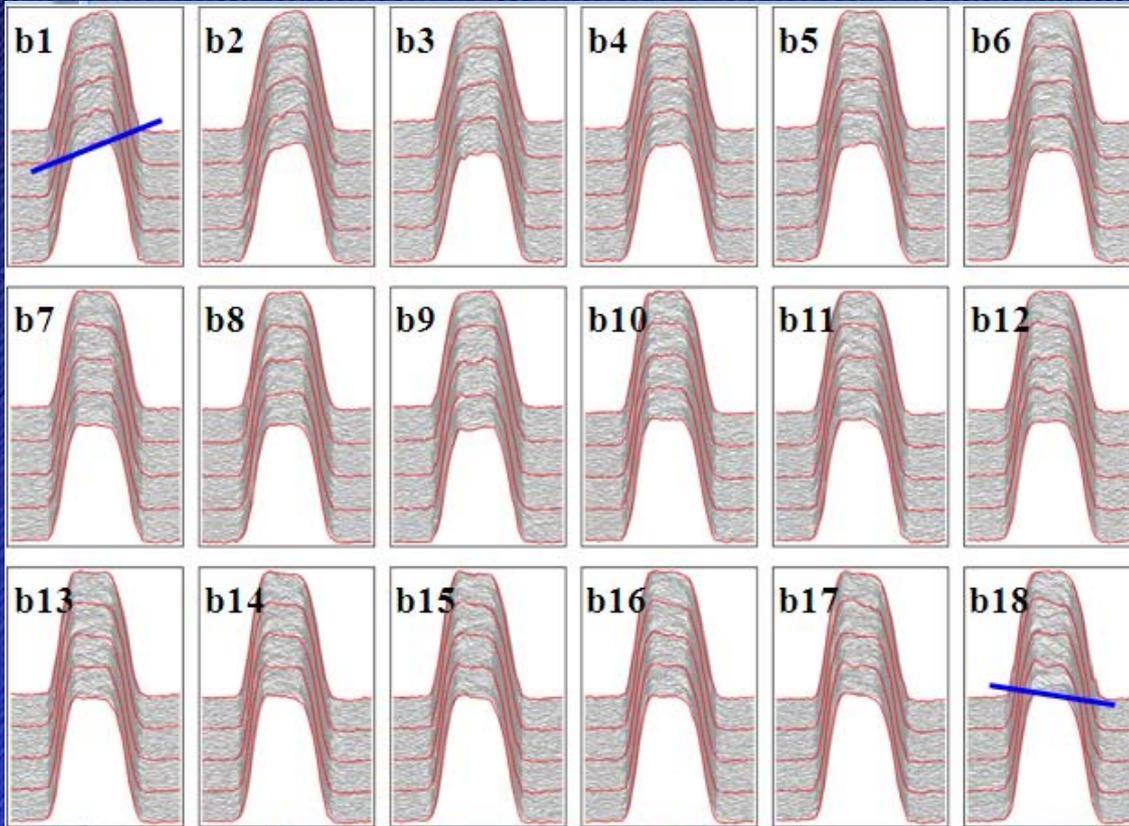
LPA progress

simulation studies and experiments on LPA beam generation & stability by Chandra Bhat (US-LARP/FNAL)

Example: Bunch Flattening of the LHC Beam at 7 TeV with 400MHz and 200MHz RF systems



LPA experiments in PS & SPS



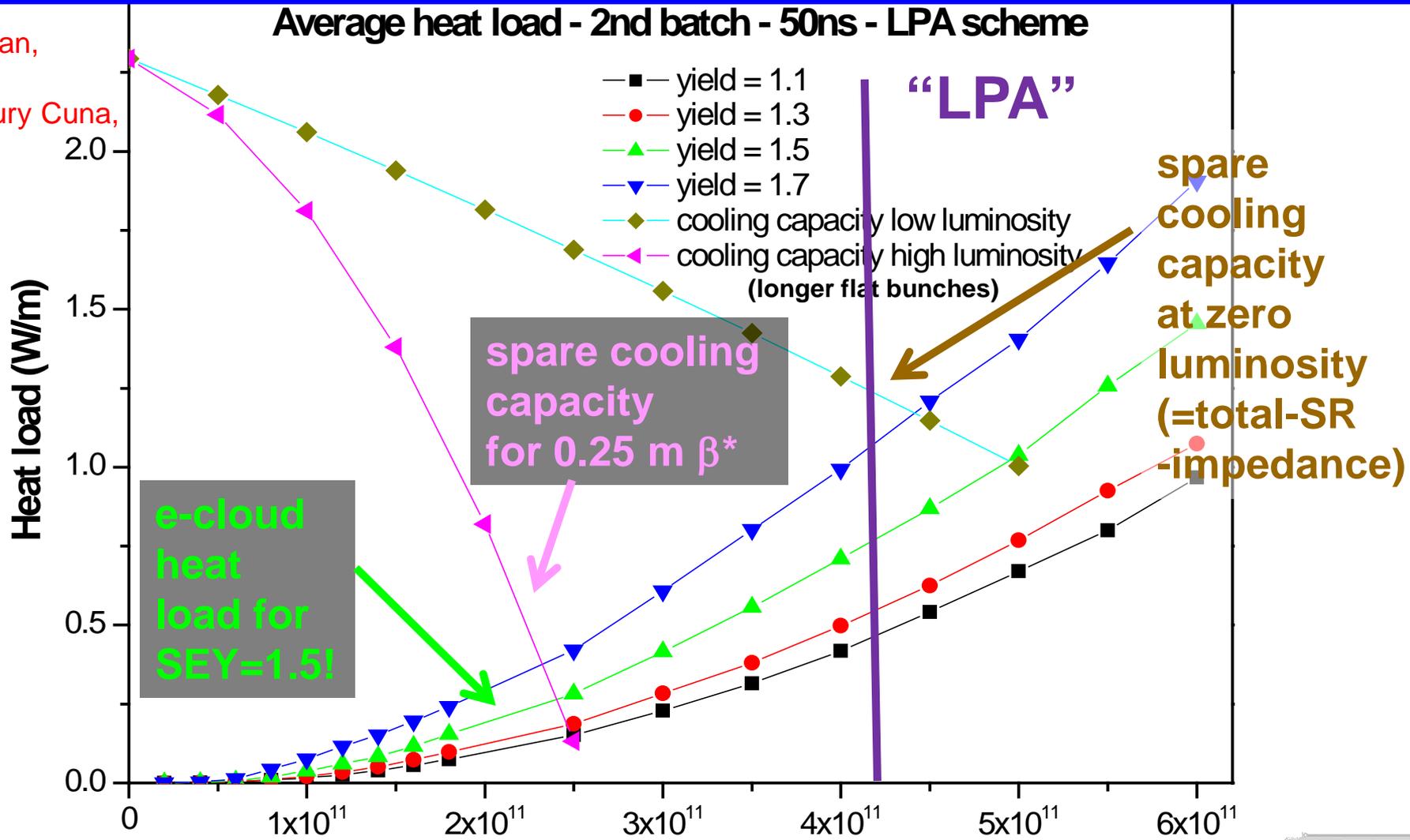
**flatness along
the PS batch**

Chandra Bhat,
Heiko Damerau,
et al.

transient beam loading compensation may be required

cooling & e- heat for 50 ns spacing

L. Taviani,
2005
H. Maury Cuna,
2009

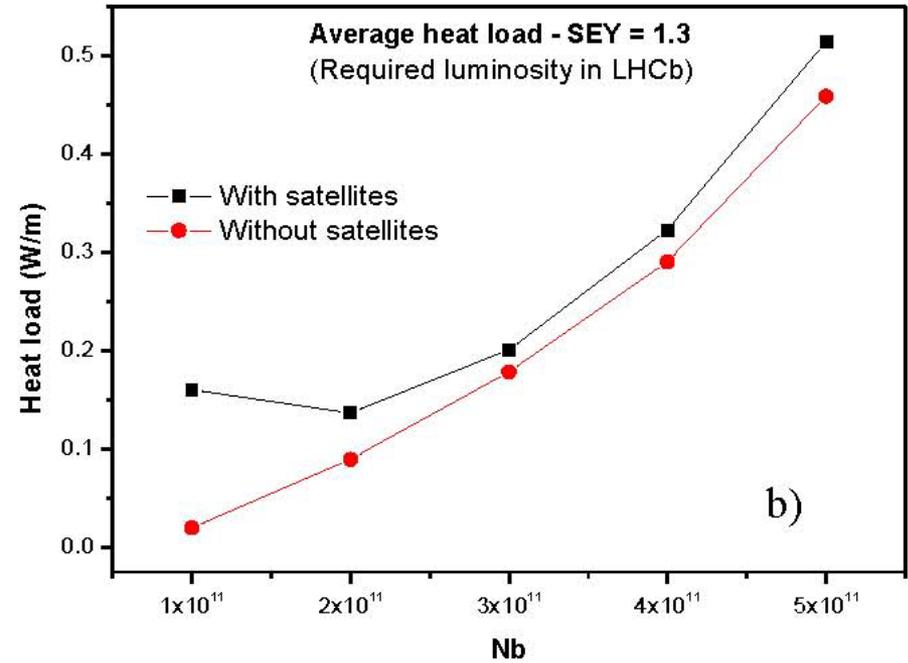
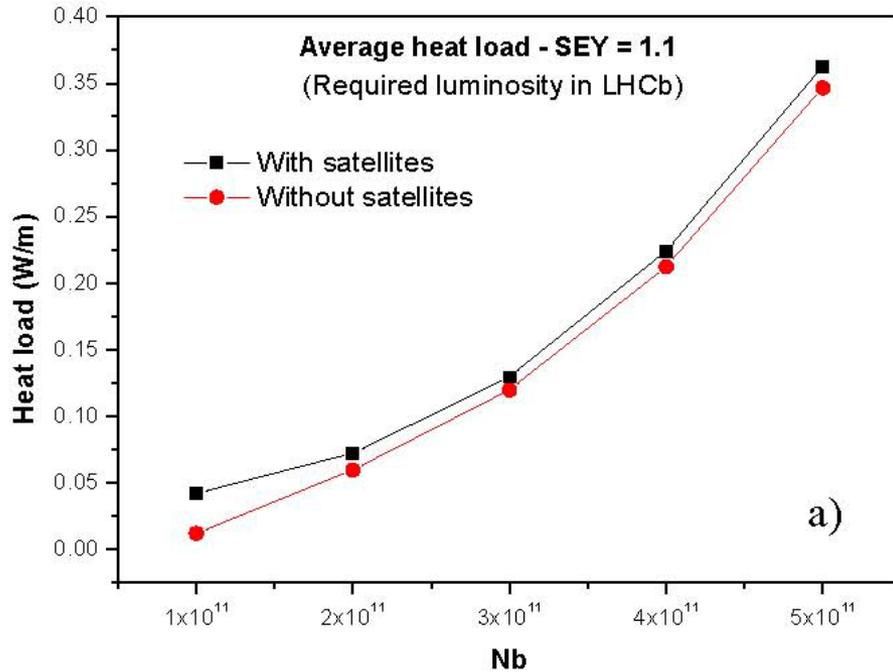


going above $N_b = 2.3 \times 10^{11}$ & ultimate luminosity requires dedicated IR cryo plants; limit then becomes $N_b \sim 5.0 \times 10^{11}$



e- heat with LHCb satellite

H. Maury Cuna, 2009



satellite intensity is varied as the inverse of main-bunch intensity to yield target luminosity of $2 \times 10^{33} \text{ cm}^{-2} \text{ s}^{-1}$ in (S)LHCb

“LHCb satellite” has small effect on 50-ns heat load

constraints - N_b range

- **beam-beam tune shift** of “head-on” collision
 - ✓ is the limit for crab crossing;
 - ✓ going beyond ultimate N_b requires large Piwinski angle or large emittance;
 - ✓ even larger crossing angle than for LR-BB may be needed in some scenarios
- **arc cooling capacity** (global & local limits)
- **collimation efficiency & machine protection**
- **injectors**

N_b constraint: collimator damage

- studied in simulations & experiments, small beam size
- **critical failure mode**: one dump kicker module pre-fires asynchronously & kicks bunches onto collimators
- **collimator damage limit in kJ/mm²**:
 - Cu: 50 kJ/mm²
 - **CFC: 5 MJ/mm² (collimators – 2 MJ/mm² tested in TT40)**
- typical location: $\sigma_r = 0.2$ mm $\rightarrow A_b = 0.13$ mm² (nominal emittance, without dilution from showers).
- stored energy & transverse energy density:
 - nominal bunch: **130 kJ \rightarrow 1.0 MJ/mm²**
 - ultimate bunch: **190 kJ \rightarrow 1.5 MJ/mm²**
 - 2 x ultimate bunch: **380 kJ \rightarrow 3.0 MJ/mm²**
- single bunch $> 5.1e11$ p exceeds damage limit of primary & secondary collimators; **damage limit depends only on total beam intensity**

constraint - beam brightness

- transverse energy density rises strongly with beam energy (γ); it also scales with number of protons (N_p^{tot}) over normalized emittance (ϵ_n):

$$\rho_E = \gamma^2 \cdot \frac{N_p^{tot}}{\epsilon_n} \cdot C \qquad C = \frac{m_p c^2}{\pi \sqrt{\beta_x \beta_y}}$$

- higher intensity or smaller emittance put similar strain on material survival!**
- “low emittance” upgrade options are no magic bullet; they solve some issues (RF, radiation, ...), but do not address damage limit
- constraint** from machine robustness:

$$\frac{N_p^{tot}}{\epsilon_n} \leq 1.3 \times 10^{20} \frac{\text{protons}}{\text{m rad}}$$

constraint – β^* range

0.55 m nominal

0.50 m ultimate

0.40 m }
0.30 m } IR “phase I” ,
0.25 m } larger aperture NbTi quad’s +...

0.22 m }
... } IR “phase II”
0.14 m } Nb₃Sn quad’s + ...
hard limit from linear chromatic correction

constraint – pile up

bunch collision rate

= #bunches/beam x revolution frequency

#events per bunch crossing

= cross section x luminosity / bunch collision rate

nominal #events/crossing in the detector

= $6 \times 10^{-26} \text{ cm}^2 \cdot 10^{34} \text{ cm}^{-2} \text{ s}^{-1} / (32 \times 10^6 \text{ s}^{-1})$

= 19

inelastic cross section



e.g. 10 times higher luminosity at same #bunches

→ ~200 events per crossing (*detector upgrade!*)

luminosity decay & lifetime

fast decay of beam intensity and luminosity (few hours)
dominated by proton burn off

$$L(t) = \frac{\hat{L}}{\left(1 + t / \tau_{eff}\right)^2}$$

with

$$\tau_{eff} = \frac{N_b n_b}{n_{IP} \hat{L} \sigma_{tot}}$$

algebraic (\neq exponential) decay!

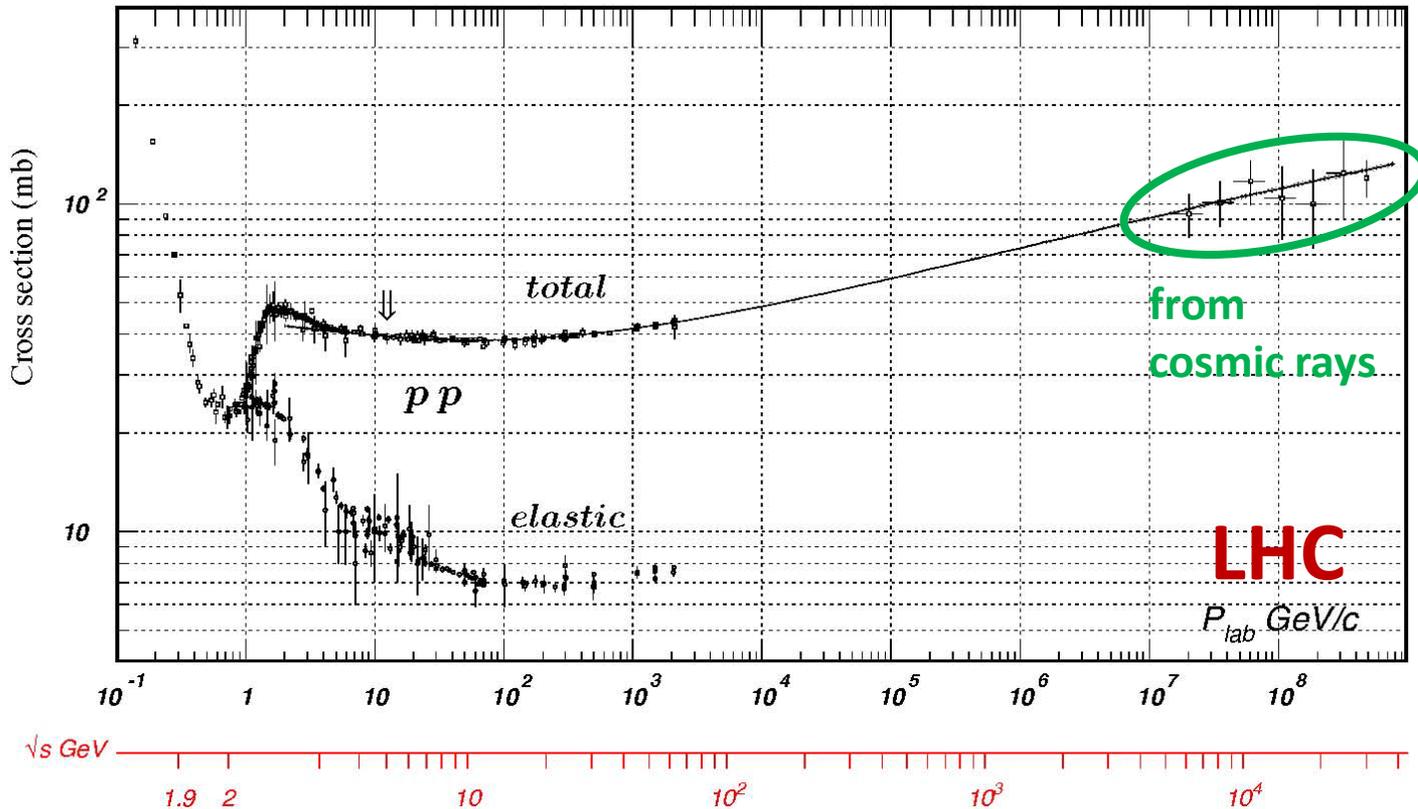
(gas scattering and IBS add negligible contributions [F.Z. ABP-RLC 23.09.05],
which are not exponential either)

$$\tau_{lumi} \propto \frac{\text{total beam intensity}}{\text{luminosity}}$$

for a given luminosity value, the luminosity lifetime
depends only on total beam current [w/o leveling]

cross sections

C. Amsler *et al.*, Physics Letters **B667**, 1 (2008)



$\sigma_{tot} \sim$
100 mbarn
 $\sim 10^{-25} \text{ cm}^2$

$\sigma_{inelastic} \sim$
60 mbarn
 $\sim 6 \times 10^{-26} \text{ cm}^2$

total cross section for LHC c.m. energy from cosmic ray experiments

example scenarios

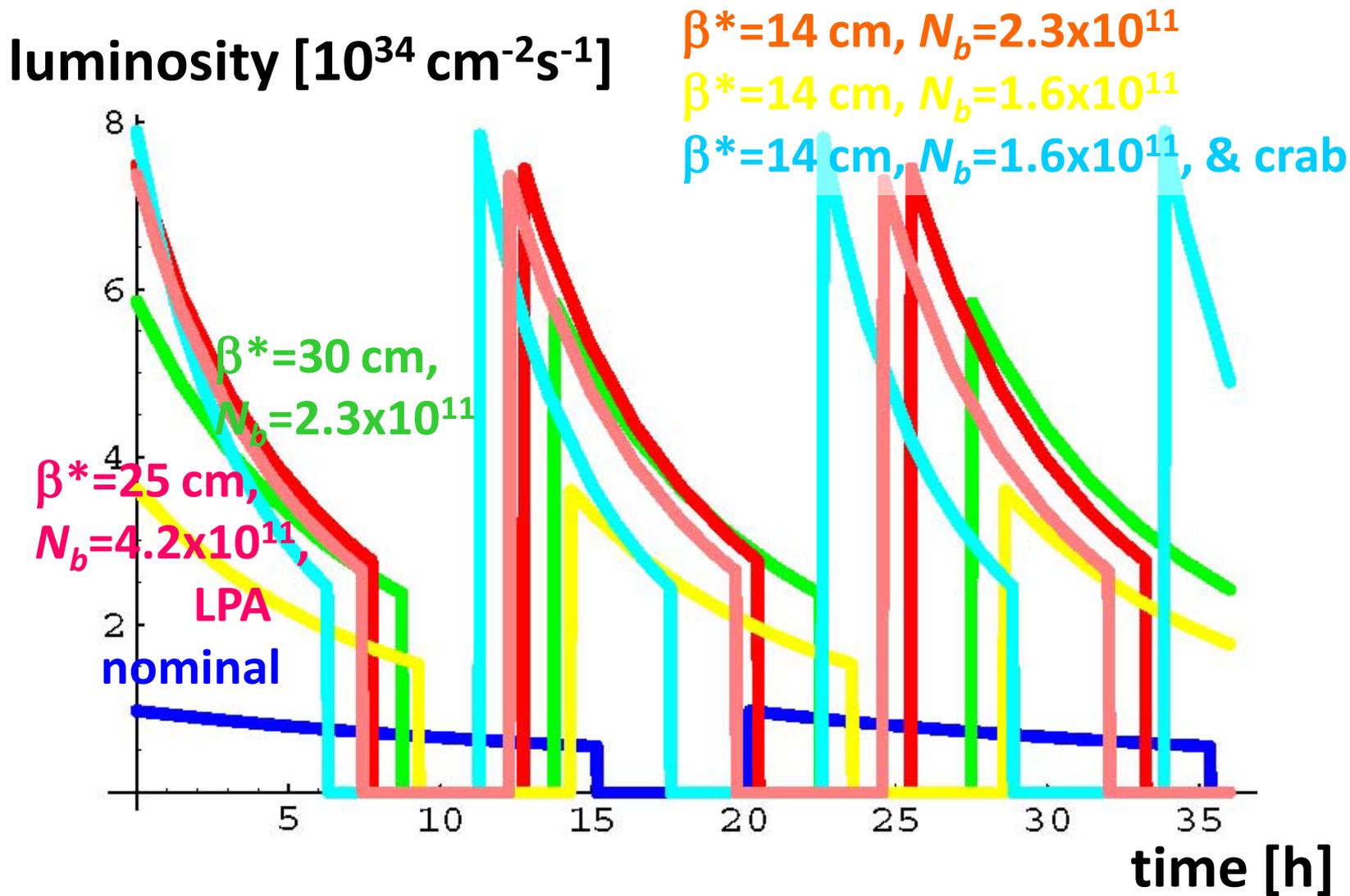
- (1) nominal, $N_b=1.15 \times 10^{11}$, $\beta^*=0.55$ m, $\theta_c=285$ μ rad
- (2) ultimate, $N_b=1.7 \times 10^{11}$, $\beta^*=0.50$ m, $\theta_c=315$ μ rad
- (3) “phase I+”, $N_b=2.3 \times 10^{11}$, $\beta^*=0.30$ m, $\theta_c=348$ μ rad
- (4) “phase I w crab”, $N_b=1.6 \times 10^{11}$, $\beta^*=0.30$ m ($\theta_c=348$ μ rad)
- (5) “phase II+”, $N_b=2.3 \times 10^{11}$, $\beta^*=0.14$ m, $\theta_c=509$ μ rad
- (6) “phase II w crab”, $N_b=1.6 \times 10^{11}$, $\beta^*=0.14$ m
($\theta_c=509$ μ rad) [also same case w/o crab]
- (7) “LPA-50”, 50 ns, $N_b=4.2 \times 10^{11}$, $\beta^*=0.25$ m, $\theta_c=381$ μ rad
- (8) “LPA-25”, 25 ns, $N_b=2.6 \times 10^{11}$, $\beta^*=0.50$ m, $\theta_c=339$ μ rad

parameter	symbol	nom.	ult.	$\beta^*=30$, cm, HI	$\beta^*=30$, cm, CC	$\beta^*=14$, cm HI	$\beta^*=14$, cm, CC	LPA – 25	LPA – 50
transverse emittance	ε [μm]	3.75	3.75	3.75	3.75	3.75	3.75	3.75	3.75
protons per bunch	N_b [10^{11}]	1.15	1.7	2.3	1.6	2.3	1.6	2.6	4.2
bunch spacing	Δt [ns]	25	25	25	25	25	25	25	50
beam current	I [A]	0.58	0.86	1.16	0.81	1.16	0.81	1.32	1.06
longitudinal profile		Gauss	Gauss	Gauss	Gauss	Gauss	Gauss	Flat	Flat
rms bunch length	σ_z [cm]	7.55	7.55	7.55	7.55	7.55	7.55	11.8	11.8
beta* at IPI&5	β^* [m]	0.55	0.5	0.30	0.30	0.14	0.14	0.50	0.25
full crossing angle	θ_c [μrad]	285	315	318	(348)	509	(509)	339	381
Piwinski parameter	$\phi = \theta_c \sigma_z / (2^* \sigma_x^*)$	0.65	0.75	1.0	0.0	2.3	0.0	2.0	2.0
tune shift	ΔQ_{tot}	0.001	0.009	0.01	0.01	0.006	0.01	0.01	0.01
peak luminosity	L [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]	1	1.5	5.9	4.0	7.5	7.9	4.0	7.4
peak events per #ing		19	21	111	76	142	150	75	280
initial lumi lifetime	T_{lum} [h]	23	15	7.7	8.8	10	11	12.4	5.3
effective luminosity ($T_{\text{turnaround}}=10$ h)	L_{eff} [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]	0.45	0.90	1.8	1.5	2.0	1.7	1.5	1.9
	$T_{\text{run,opt}}$ [h]	21.5	17.2	1.4	2.5	11.0	8.9	16.0	10.5
effective luminosity ($T_{\text{turnaround}}=2$ h)	L_{eff} [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]	0.67	1.1	3.2	2.2	3.8	3.5	2.4	3.6
	$T_{\text{run,opt}}$ [h]	15.6	7.7	5.5	5.6	4.9	4.0	7.2	4.7
e-c heat SEY=1.3	P_{e-c} [W/m]	0.4	0.6	1.3	0.7	1.3	0.7	1.4	0.8
SR heat 4.6-20 K	P_{SR} [W/m]	0.17	0.25	0.34	0.24	0.34	0.24	0.38	0.31
image current heat	P_{ic} [W/m]	0.15	0.33	0.60	0.29	0.60	0.29	0.39	0.51
gas-s. 100 h τ_b	P_{gas} [W/m]	0.04	0.06	0.08	0.05	0.08	0.05	0.09	0.07
luminous region	σ_1 [cm]	4.5	4.3	3.7	5.3	2.2	5.3	5.2	3.8
annual luminosity	L_{int} [fb^{-1}]	57	116	245	169	286	253	198	274

parameter highlights

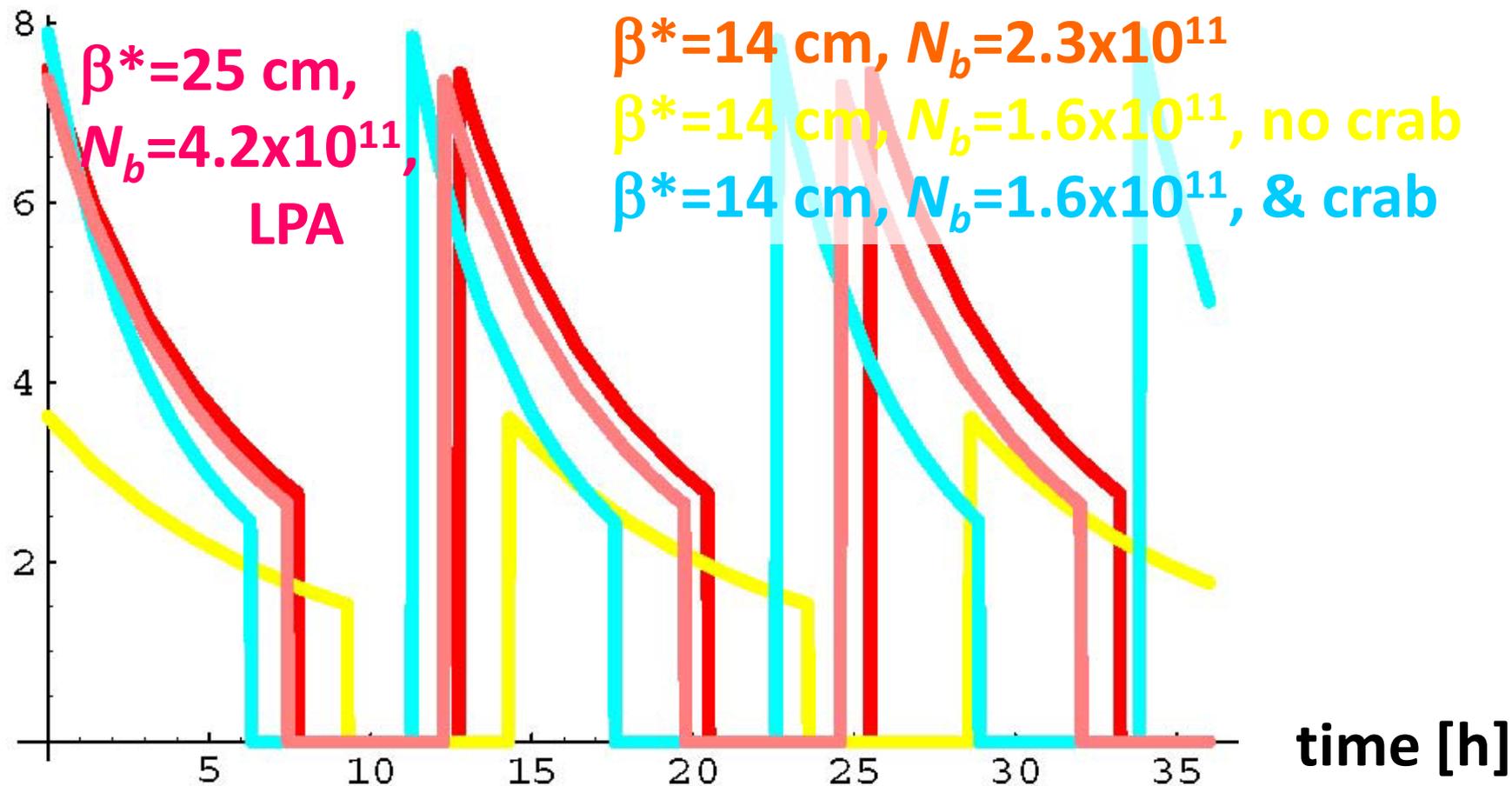
parameter	symbol	nom.	ult.	$\beta^*=30$	$\beta^*=30$ (crab)	$\beta^*=14$	$\beta^*=14$ (crab)	LPA(50 ns, flat)
ppb	$N_b [10^{11}]$	1.15	1.7	2.3	1.6	2.3	1.6	4.2
beta* at IP1&5	$\beta^* [m]$	0.55	0.5	0.30	0.30	0.14	0.14	0.25
Piwinski angle		0.65	0.75	1.1	0.0	2.3	0.0	2.0
tune shift	ΔQ_{tot}	0.009	0.009	0.01	0.01	0.006	0.01	0.01
peak luminosity	$L [10^{34}$ $cm^{-2}s^{-1}]$	1	2.3	5.9	4.0	7.5	7.9	7.4
peak evt's / #ing		19	44	111	76	142	150	280
lumi lifetime	$\tau_L [h]$	23	15	7.7	7.8	6.0	4.0	5.3
average ($T_{turnaround}=5$ h)	$L_{eff} [10^{34}$ $cm^{-2}s^{-1}]$	0.55	1.12	2.4	1.6	2.8	2.4	2.6
	$T_{run,opt} [h]$	15.2	12.2	8.7	8.8	7.7	6.3	7.5
annual luminosity (200 days, 60% availability)	$L_{int} [fb^{-1}]$	57	116	245	168	286	253	274

luminosity evolution - examples



luminosity evolution – selected cases

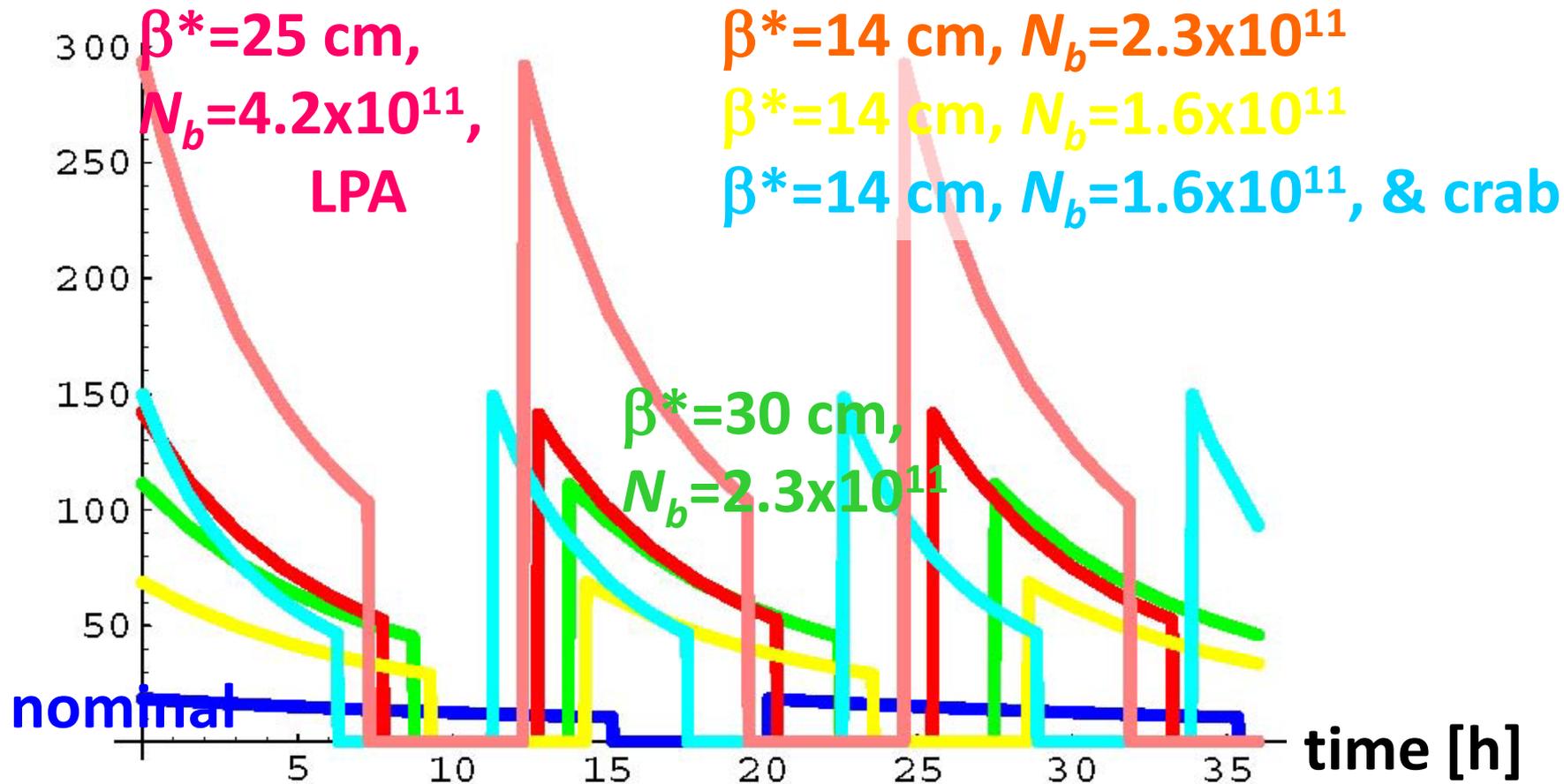
luminosity [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]



$\beta^*=14 \text{ cm} \ \& \ N_b=2.3 \times 10^{11}$ has very similar performance to $\beta^*=14 \text{ cm}, \ \& \ N_b \sim 1.6 \times 10^{11}$ and crab, and to $\beta^*=25 \text{ cm} \ \& \ N_b=4.2 \times 10^{11}$ & 50 ns spacing

events/crossing evolution

#events/crossing



all scenarios give peak #events/#ing $\sim 100-150$,
except for LPA ~ 300

luminosity leveling

changing θ_c , β^* or σ_z during the store in order to
→ **reduce event pile up & IR peak power deposition**
→ **maximize integrated luminosity**

leveling with crossing angle has two advantages:

increased average luminosity, operational simplicity

natural option for early separation or crab cavities,

leveling may first be tested in LHC heavy-ion collisions

two leveling strategies:

(1) constant luminosity

(2) constant beam-beam tune shift

optimum run time & av. luminosity

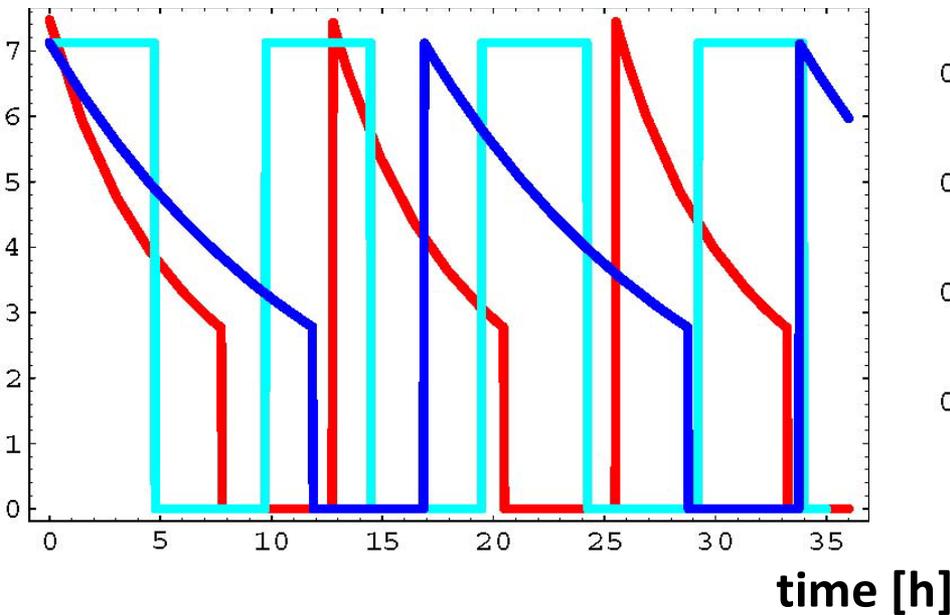
	w/o leveling	$L=\text{const}$	$\Delta Q_{\text{bb}}=\text{const}$
luminosity evolution	$L(t) = \frac{\hat{L}}{(1+t/\tau_{\text{eff}})^2}$	$L = L_0 \approx \text{const}$	$L(t) = \hat{L} \exp(-t/\tau_{\text{eff}})$
beam current evolution	$N(t) = \frac{N_0}{(1+t/\tau_{\text{eff}})}$	$N = N_0 - \frac{N_0}{\tau_{\text{eff}}} t$	$N(t) = N(0) \exp(-t/\tau_{\text{eff}})$
optimum run time	$T_{\text{run}} = \sqrt{\tau_{\text{eff}} T_{\text{ta}}}$	$T_{\text{run}} = \frac{\Delta N_{\text{max}} \tau_{\text{eff}}}{N_0}$	$T_{\text{run}} = \tau_{\text{eff}}$ $\min \left[\ln \left(\sqrt{1 + \phi_{\text{piw}}(0)^2} \right), \right.$ $\left. \ln \left((T_{\text{ta}} + T_{\text{run}} + \tau_{\text{eff}}) / \tau_{\text{eff}} \right) \right]$
average luminosity	$L_{\text{ave}} = \hat{L} \frac{\tau_{\text{eff}}}{(\tau_{\text{eff}}^{1/2} + T_{\text{ta}}^{1/2})^2}$	$L_{\text{ave}} = \frac{L_0}{1 + \frac{L_0 \sigma_{\text{tot}} n_{\text{IP}} T_{\text{ta}}}{\Delta N_{\text{max}} n_b}}$	$L_{\text{ave}} = \frac{\tau_{\text{eff}}}{T_{\text{ta}} + T_{\text{run}}} \left(1 - e^{-T_{\text{run}}/\tau_{\text{eff}}} \right)$

leveling 2 \rightarrow exponential L decay, w decay time τ_{eff} (not $\tau_{\text{eff}}/2$)

leveling – example evolution

$$\beta^* = 14 \text{ cm}, N_b = 2.3 \times 10^{11}, T_{ta} = 5 \text{ h}$$

luminosity [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]

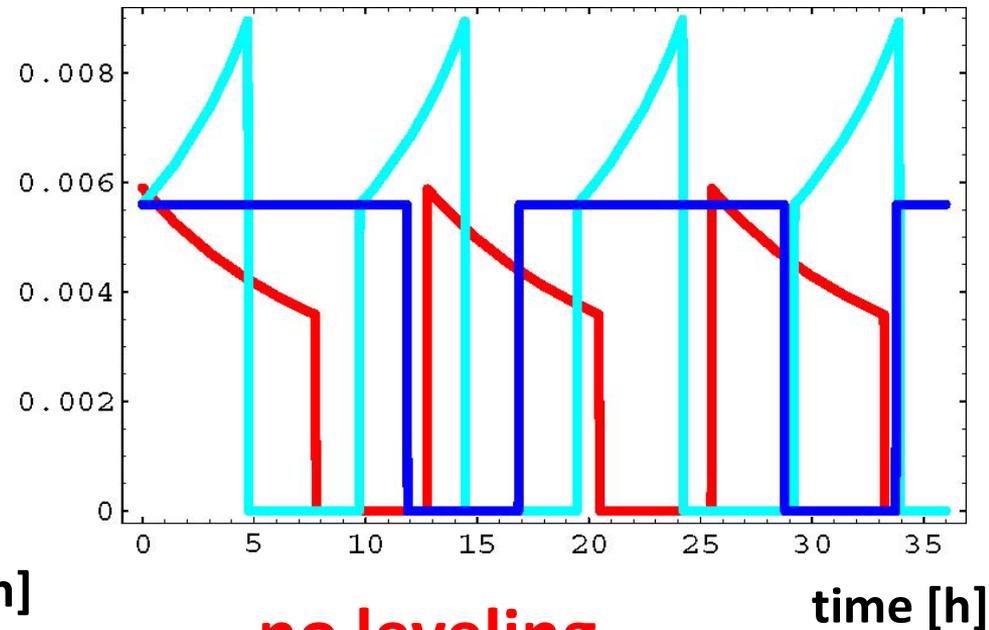


no leveling

$\Delta Q = \text{const}$

$L = \text{const}$

$|\Delta Q|$



no leveling

$\Delta Q = \text{const}$

$L = \text{const}$

leveling – example numbers

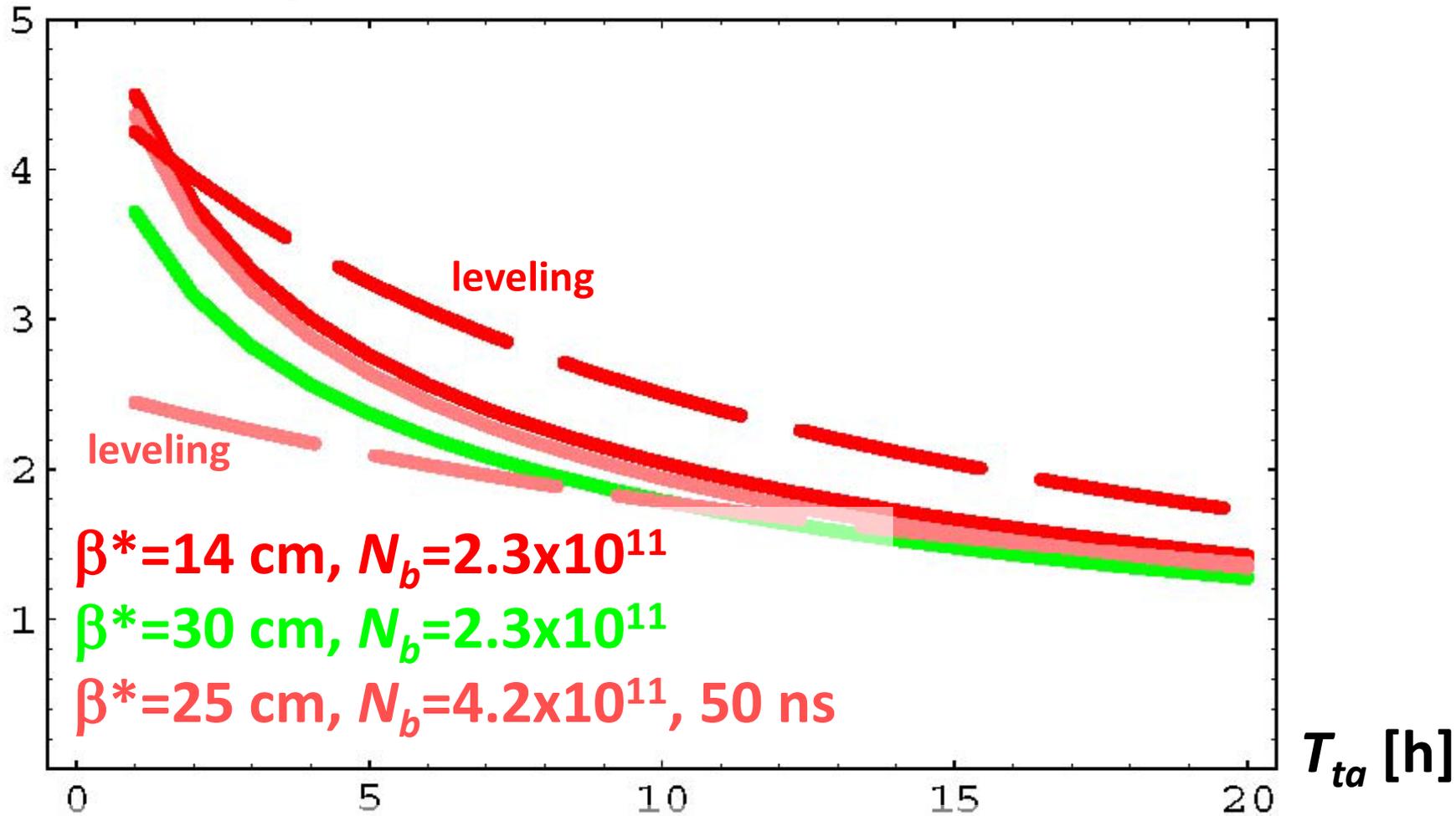
	$\beta^*=14$ cm, 25 ns spacing, $T_{ta}=5$ h			
	no leveling	$L=\text{const}$	$\Delta Q_{bb}=\text{const}$	
$N_b(0)$ [10^{11}]	2.3	2.3	2.3	2.3
$L(0)$ [$10^{34}\text{cm}^{-2}\text{s}^{-1}$]	7.5	7.1	12.3	7.1
$ \Delta Q_{bb}(0) $	0.0059	0.0056	0.01	0.0056
$ \Delta Q_{bb}(T_{run}) $	0.0036	0.0090	0.01	0.0056
$\theta_c(0)$ [μrad]	509	539	239	539
run time T_{run} [h]	7.74	4.74	2.72	11.9
$\langle L \rangle$ [$10^{34}\text{cm}^{-2}\text{s}^{-1}$]	2.8	3.5	3.6	3.2
events/#ing (0)	149	135	234	35

leveling – other example numbers

	$\beta^*=25$ cm, 50 ns spac., “LPA” $T_{ta}=5$ h		
	no leveling	$L=\text{const}$	$\Delta Q_{bb}=\text{const}$
$N_b(0)$ [10^{11}]	4.2	4.2	4.2
$L(0)$ [$10^{34}\text{cm}^{-2}\text{s}^{-1}$]	7.4	4.5	4.5
$ \Delta Q_{bb}(0) $	0.010	0.0056	0.0056
$ \Delta Q_{bb}(T_{run}) $	0.006	0.010	0.0056
$\theta_c(0)$ [μrad]	231	672	672
run time T_{run} [h]	7.45	6.0	23.2
$\langle L \rangle$ [$10^{34}\text{cm}^{-2}\text{s}^{-1}$]	2.6	2.5	2.4
events/#ing (0)	280	172	172

$\langle L \rangle$ vs. turnaround time

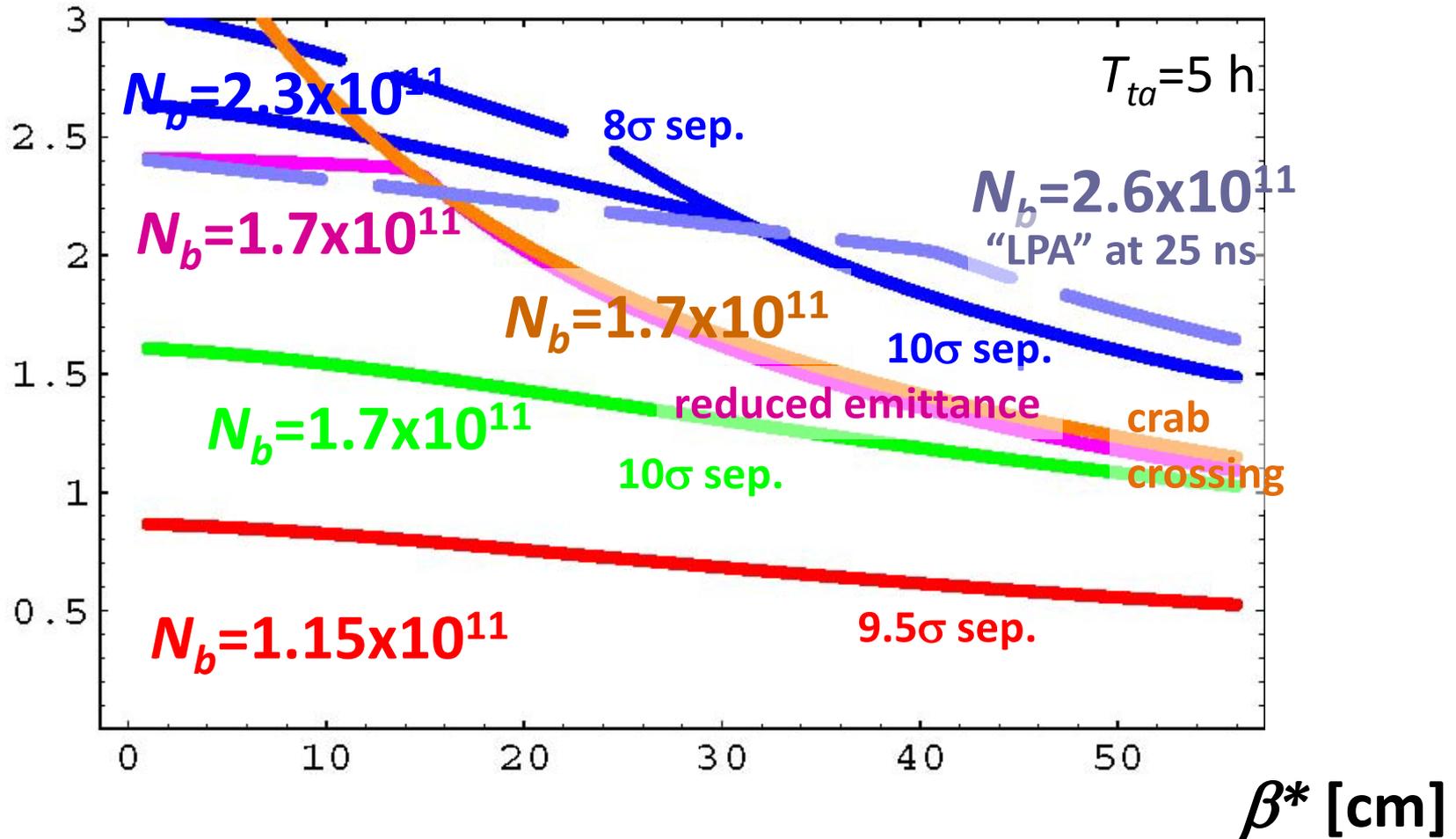
$\langle L \rangle$ [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]



reducing T_{ta} from 10 to 2 h increases $\langle L \rangle$ about 2x,
similar average luminosity for all 3 scenarios

$\langle L \rangle$ vs. β^* - the KEY PLOT

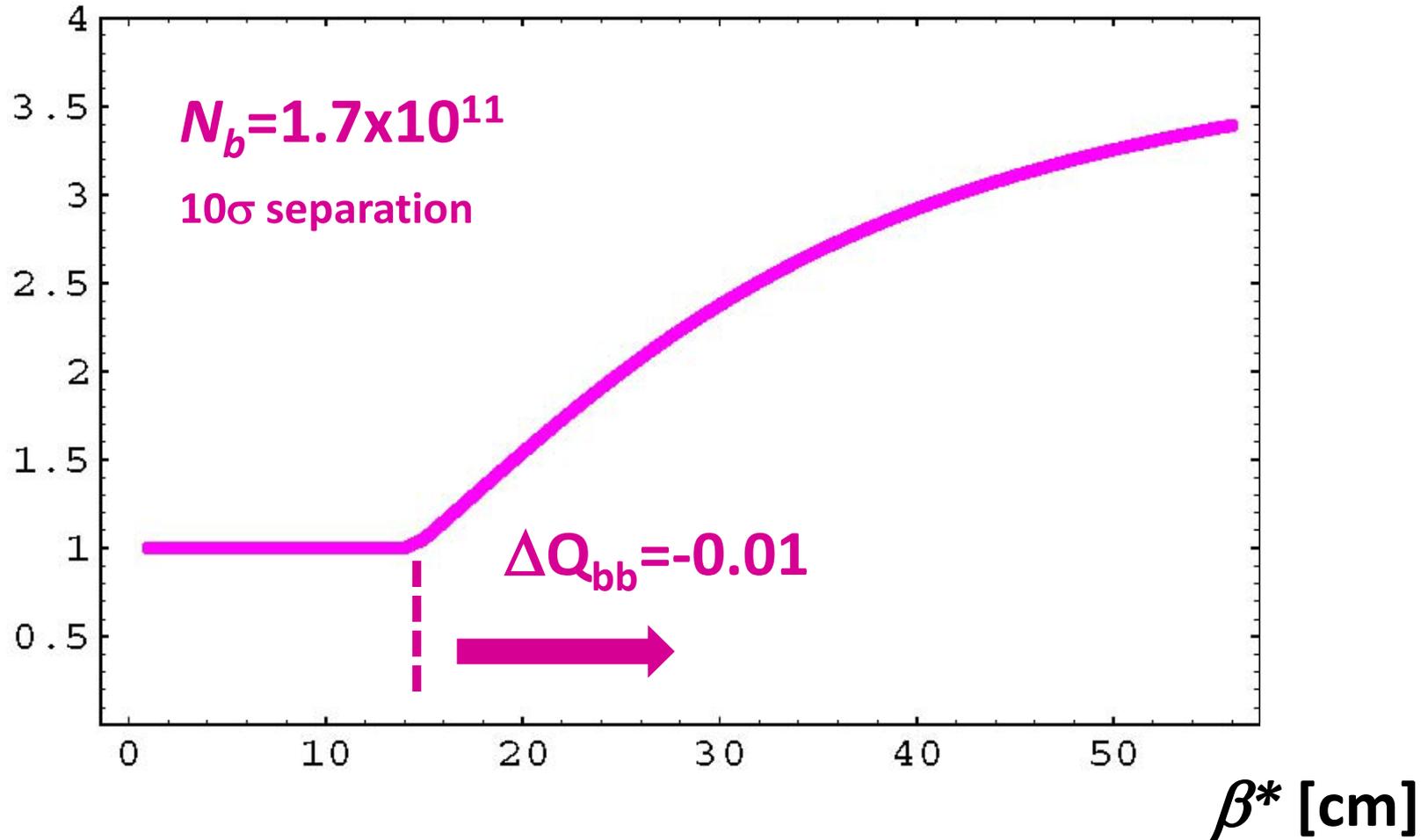
$\langle L \rangle$ [$10^{34} \text{ cm}^{-2} \text{ s}^{-1}$]



beam intensity is much more important than β^* , reducing β^* only helps with crab cavities or with smaller emittance

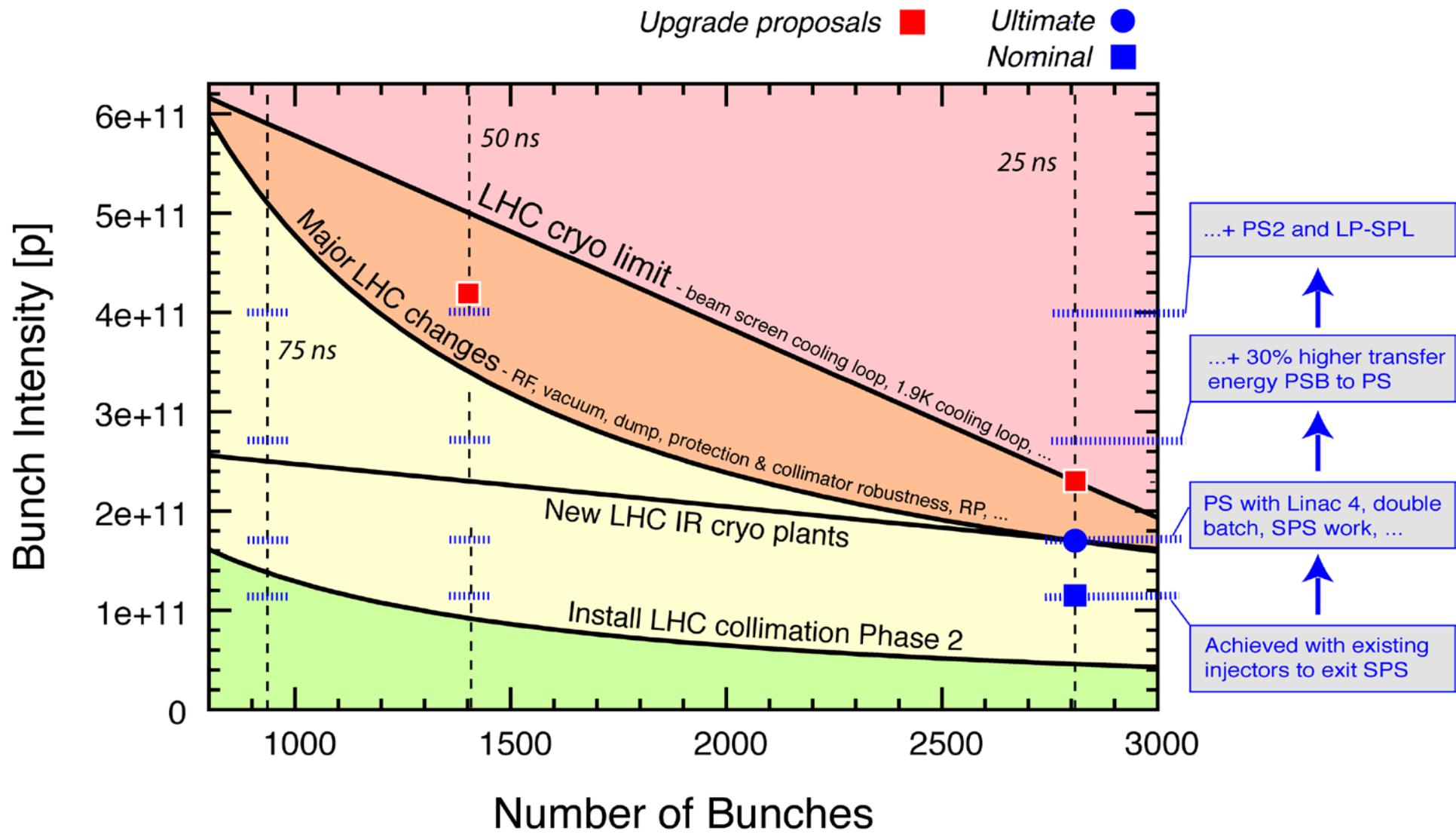
ϵ vs. β^* - for low-emittance scheme

$\gamma\epsilon$ [μm]



emittance for the low-emittance scheme determined by ΔQ

LHC intensity limits at 7 TeV



conclusions

- several upgrade scenarios w. 25 or 50-ns spacing
- annual luminosities of 150-300 fb⁻¹
- collimation phase 2 essential
- beyond ultimate: separate cryoplants for IR1, 5 & 4
- maximum $N_b \sim 2.3 \times 10^{11}$ at 25 ns, $\sim 5.0 \times 10^{11}$ at 50 ns
limited by arc beam-screen cooling capacity
- T_{ta} - 10 → 2 h: 2x higher $\langle L \rangle$
- β^* : factor 2 reduction → 10-20% higher $\langle L \rangle$, unless accompanied by crab cavities or smaller ε
- N_b : factor 2 increase → 3 times higher $\langle L \rangle$!
- crab crossing: 10-100% higher $\langle L \rangle$; crab cavities also provide easy leveling & increase flexibility

more conclusions

- leveling with (effective) crossing angle:
 - 1.5-3 x higher T_{run} , → 40% lower peak pile up
 - (or) increase $\langle L \rangle$ by ~15%
- present luminosity optimization assumes collisions in two IPs, LHCb collisions compatible with 50-ns spacing by adding less-intense satellite bunches
- recommended **R&D focus**:
 - understanding and **mitigating intensity limits**
 - **minimization of turnaround time** (3 h → ~1 h?)
 - new **interaction-region design with (much) smaller β^* together with crab cavities and/or smaller-emittance beams**

questions

- **how much event pile up is acceptable?**
 - is there a clear upper limit and which?
- is #events per crossing the relevant number, or e.g. #events per 50 ns?
 - or in other words, **is pile up limit / crossing the same for 25-ns and 50-ns spacing?**
- is there an **official policy or guideline for LHCb and ALICE running at the time of SLHC?**; will the 4 experiments always run together? present upgrade scenarios are optimized for high luminosity in two IPs; additional collisions will contribute to ΔQ_{bb}

thank you for your attention!

02.04.2010 - CENTRAL STUDIO'S UTRECHT

LUMINOSITY

BEFORE THE ENERGY